

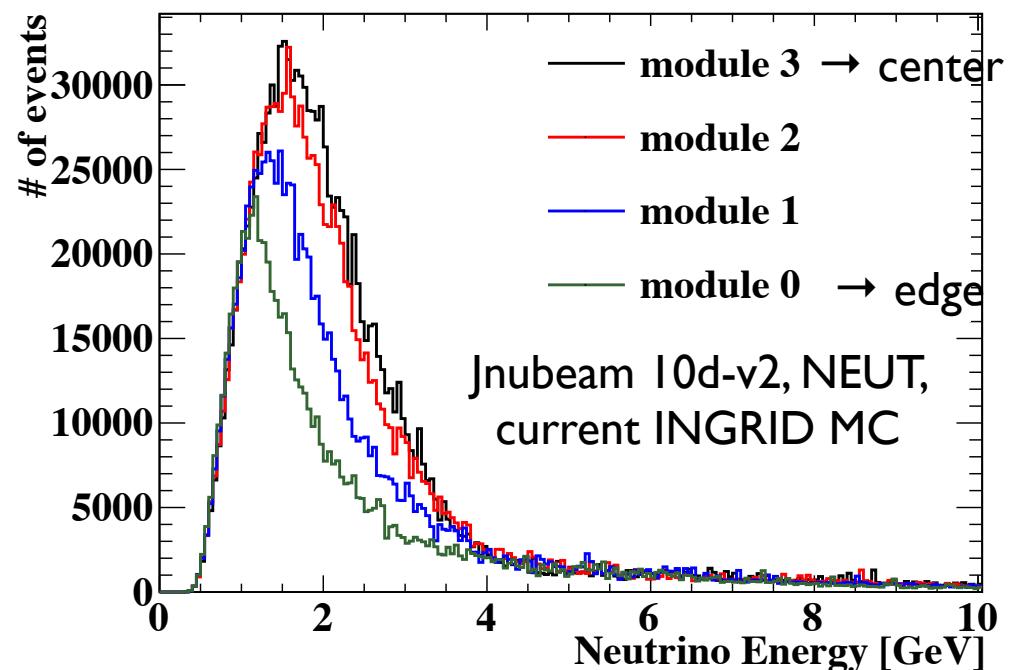
# Possibility of CC inclusive measurement w/ INGRID

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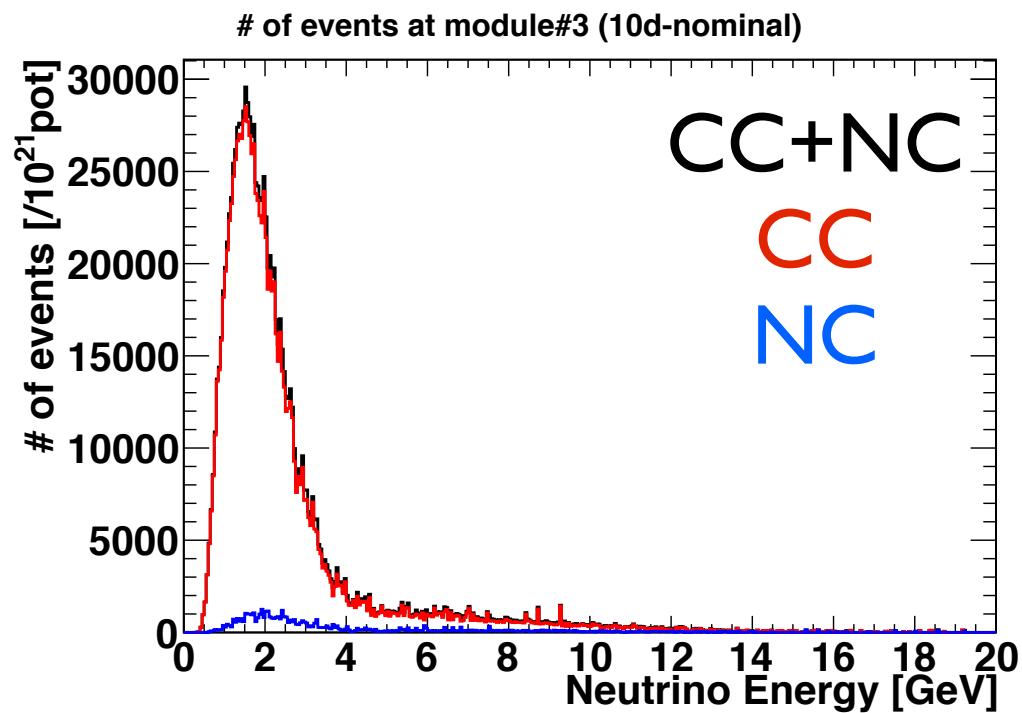
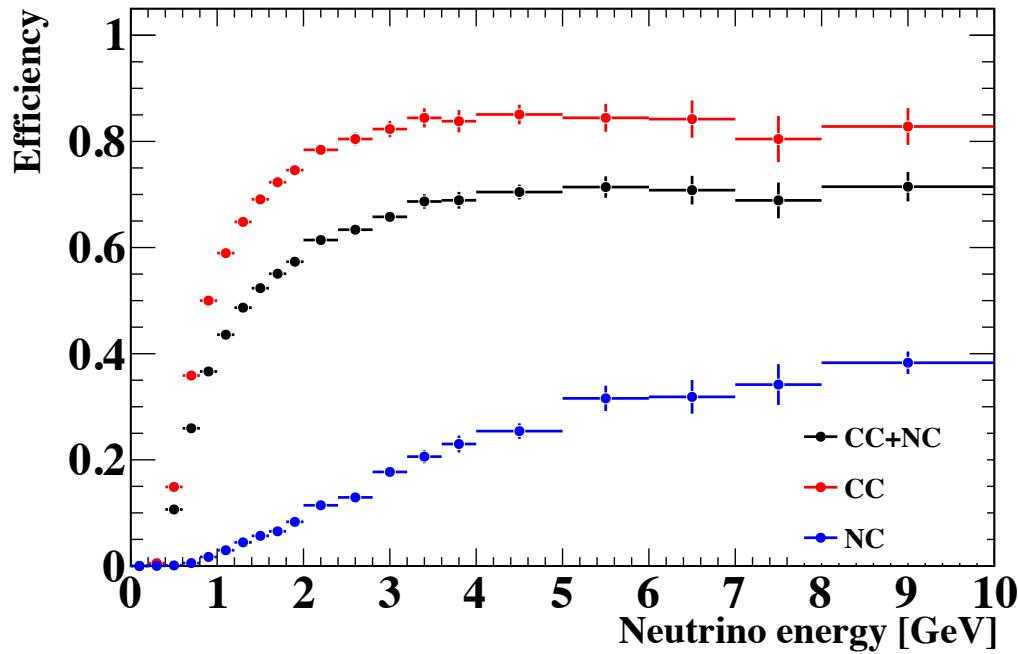
# Motivation

- Neutrino energy distribution is expected to be different at each module.
  - Because each modules is set with different off-axis angle.
- Flux at each module is almost same for  $E_{\nu} < 1 \text{ GeV}$  and  $E_{\nu} > 4 \text{ GeV}$ , but different at  $1 < E_{\nu} < 4 \text{ GeV}$
- I try to measurement the inclusive neutrino cross-section at  $\bar{E}_{\nu} \sim 4 \text{ GeV}$  (hopefully  $\sim 3 \text{ GeV}$ ) by using the flux difference among modules.

$\nu\mu$  true energy distribution after neutrino selection



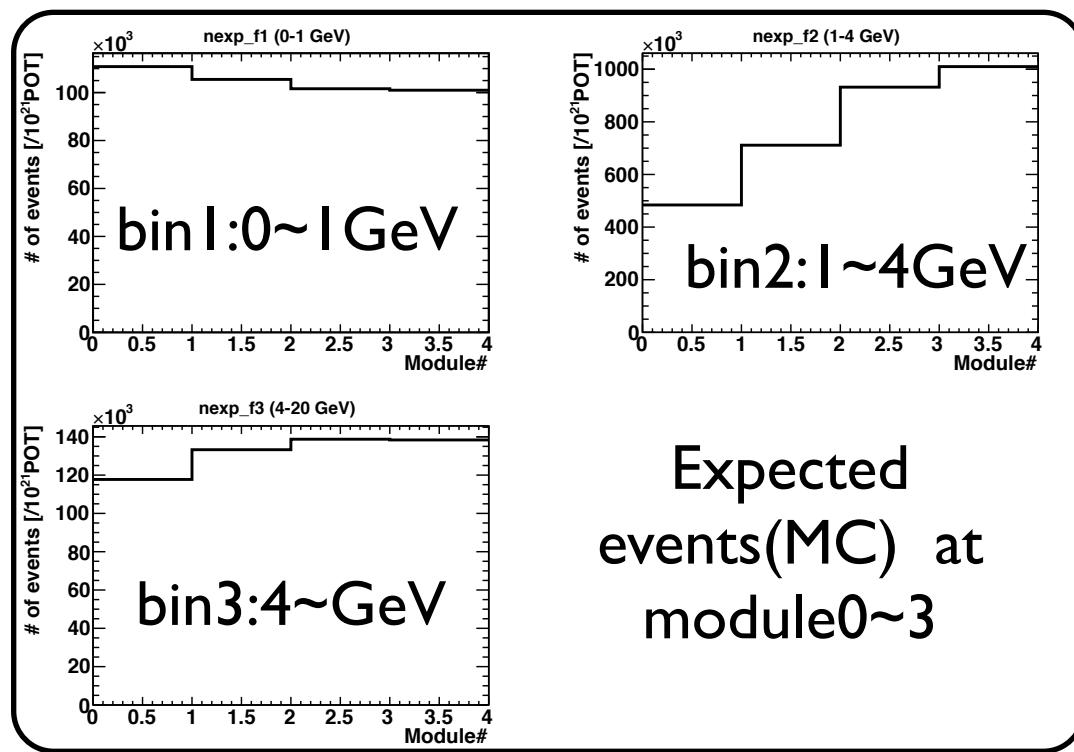
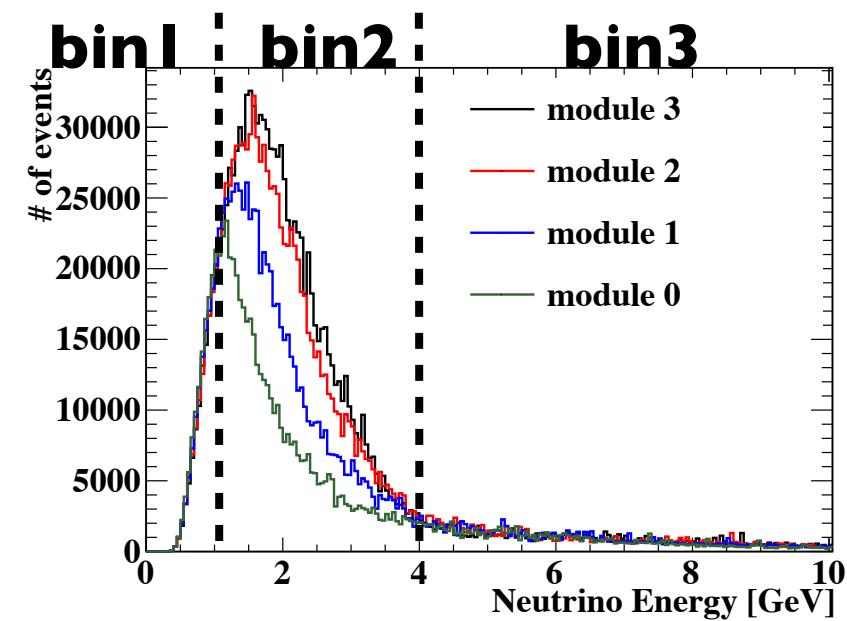
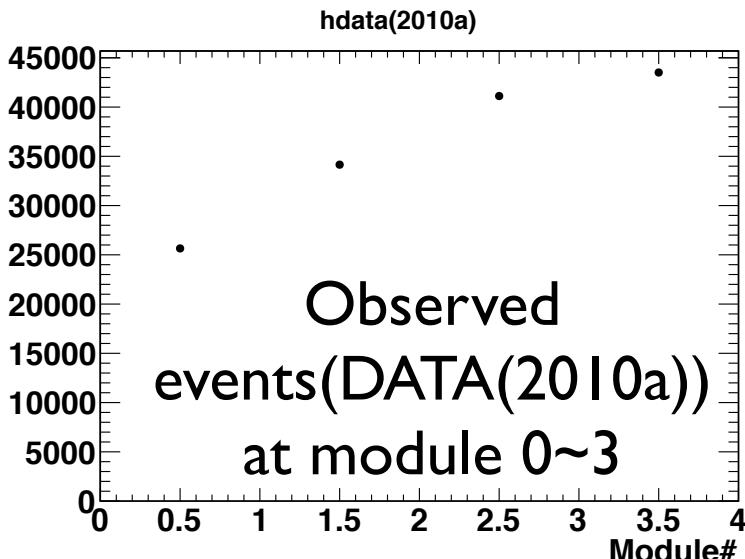
# INGRID Efficiency to neutrino interaction in FV



- More than 90% of remained events after current neutrino selection is events of CC interaction mode.
  - High CC interaction purity.

# Motivation

- Divide true neutrino energy into some bins
- Cross section is obtained as the normalization factor( $f_i$ ) at each bin( $i$ ) common to all modules.
- Get  $f_i$  which reproduce Data/MC for all modules.
- Measured Cross section = Cross section(MC)  $\times (1+f_i)$



# Fit function

Maximum the following likelihood:

$$L = \prod_{\text{module}} \text{Gauss}(N_{\text{mod}}^{\text{obs}} | N_{\text{mod}}^{\text{exp}}, \sqrt(N_{\text{mod}}^{\text{exp}})) \quad (\text{stat. error}) \\ \times \frac{1}{2\pi^{n/2} \sqrt(\det \Sigma_{\text{sys}})} \exp(-\frac{1}{2}(f_1, f_2, f_3, \dots) \Sigma_{\text{sys}}^{-1} (f_1, f_2, f_3, \dots)^T) \quad (\text{syst. error})$$

$$N_{\text{mod}}^{\text{exp}} = \Sigma_{\text{bins}} \int_{\text{bin}\#i} \phi(E) \sigma(E) \epsilon(E) \times (1 + f_i) dE$$

- $(f_1, f_2, f_3, \dots)$  are fitting parameters (normalization at each energy bins)
- At first step, consider only stat. error term
  - Fitting parameters is constrained :  $1 + f_i > 0$ .
  - I tried some division of energy bins.

# Fitting demonstration

- Check whether fitting is valid or not by using Fake MC ( $N^{\text{exp,fake}}$ ) as  $N^{\text{obs}}$ .
  - Fake MC is changed from MC expectation by design as like the following formula:

$$N_{\text{mod}}^{\text{exp,fake}} = \sum_{\text{bins}} \int_{\text{bin}\#i} \phi_{\text{mod}}(E) \sigma(E) \epsilon(E) (1 + \delta_i) dE$$

$$\delta_i = 0.1 * i \quad \rightarrow \text{Use this } N_{\text{mod}}^{\text{exp,fake}} \text{ as } N_{\text{mod}}^{\text{obs}}$$

- Check if obtained parameters ( $f_i$ ) by fitting are consistent to given  $\delta_i$ , my fitting analysis is valid.
- Normalization parameters I want to measure have few ten % error at this time with only stat. error, this analysis method needs to be change.

# Analysis configuration

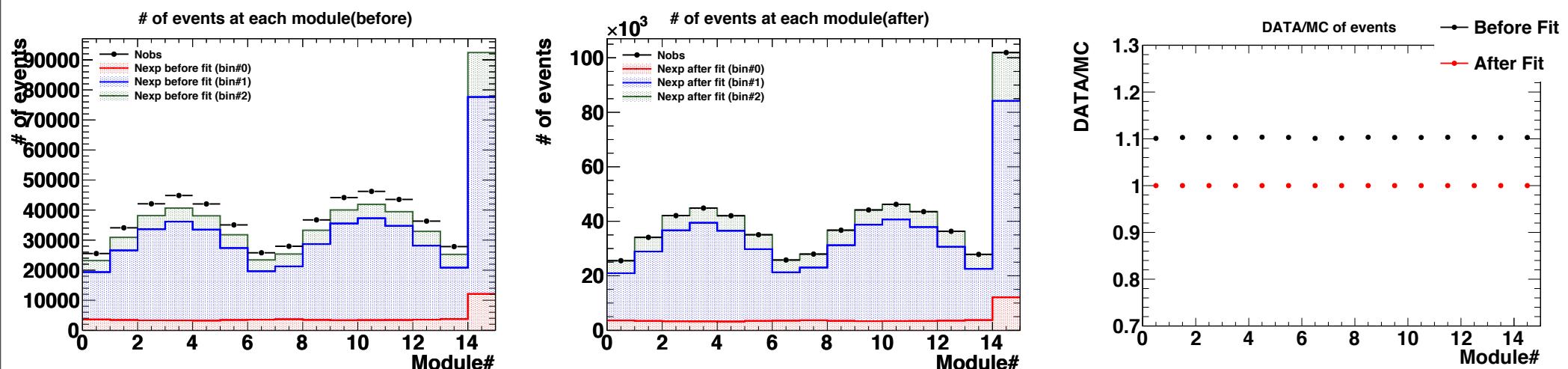
- Used modules for analysis : #0~3, #7~10, #14(shoulder module)
- MC configuration
  - Flux : Jhubeam 10d-v2 (by weighting based on energy)
  - Neutrino interaction : NEUT 5.0.6
  - Detector MC : current INGRID MC (not nd280 software)
- MC stat. is normalized according to the followings:
  - #0~3, #7~10 : RunI p.o.t. (update RunI & RunII soon)
  - #14 : RunII p.o.t.
- Try two cases for binning as first trial
  - Case 1 :  $<1\text{GeV}$ (bin#1),  $1\sim4\text{GeV}$ (bin#2),  $>4\text{GeV}$ (bin#3)
  - Case 2 :  $<1\text{GeV}$ (bin#1),  $1\sim3\text{GeV}$ (bin#2),  $>3\text{GeV}$ (bin#3)
- In both case, cross-section in 2nd bin is expected to be measured and there would not be sensitivity to dissolve 1st and 3rd bin because of the flux difference of among modules.

# Case I

NAME	VALUE	ERROR
f1	9.99417e-01	3.51683e-01
f2	1.09996e+00	1.87102e-02
f3	1.20058e+00	3.70212e-01

```
==== INPUT Fake MC ====
bin#:0 = 1
bin#:1 = 1.1
bin#:2 = 1.2
```

- Equal to input Fake MC.
- Fitting error of parameter I want to measure is  $\sim 2\%$   $\rightarrow$  Good



$\rightarrow$  Data(Fake MC) / MC become closer to 1.

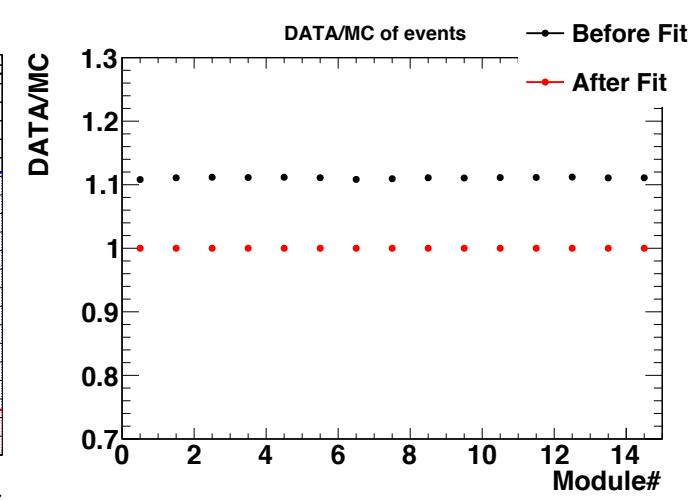
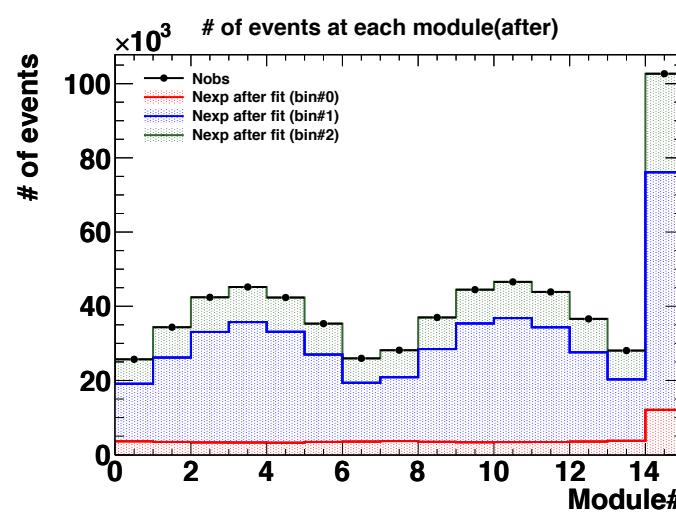
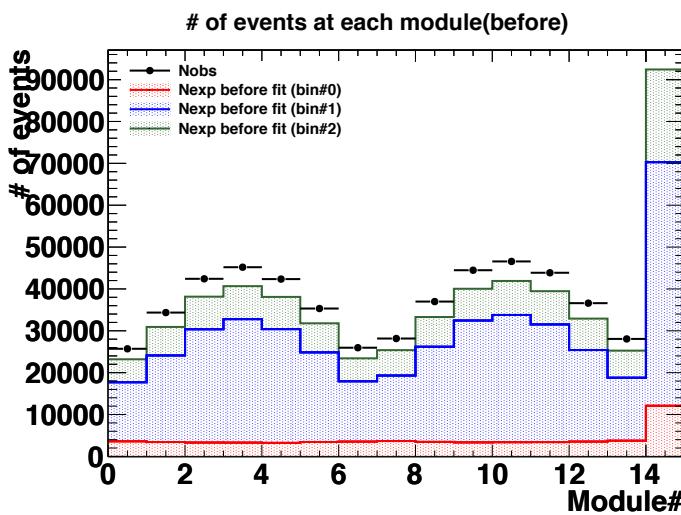
This fitting seems to be valid.

# Case 2

NAME	VALUE	ERROR
f1	9.99515e-01	2.66557e-01
f2	1.09992e+00	3.96428e-02
f3	1.20044e+00	2.52056e-01

```
==== INPUT Fake MC ====
bin#:0 = 1
bin#:1 = 1.1
bin#:2 = 1.2
```

- Equal to input Fake MC.
- Fitting error of parameter I want to measure is  $\sim 4\%$   $\rightarrow$  More than first, but small.



This fitting seems to be valid.

# List of systematic uncertainty

- Detector systematic uncertainty
    - Otani-san, Christophe, other people have studied. Need to consider treatment of uncertainty in my analysis.
  - Flux uncertainty
    - Due to hadron production, proton beam parameter, off-axis, etc.
  - Uncertainty of GEANT4 physics models
  - Uncertainty related neutrino interaction
    - Events in interacted in scintillator is expected to account for 5% of in iron. I would treat the interaction in scintillator as a uncertainty.
    - Use the current uncertainty studied by NIWG.
- ➡ At first step, I consider about well-known detector systematic

# Detector systematic

- Two kinds of error : detector common / dependent.
- At current study, detector systematic errors are considered as common (according to TN-040)).
  - Deference of each module is estimated to be small against error size.
- At first step, I consider this error as common in all modules. This part of likelihood can be defined as the followings:

$$L^{det. syst.} = \prod_{module} Gauss(N_{mod}^{exp} | N_{0,mod}^{exp}, \sigma_{N_{0,mod}^{exp}}^{det.})$$

$$N_{0,mod}^{exp} = \sum_{bins} \int_{bin\#i} \phi(E) \sigma(E) \epsilon(E) dE \quad (MC \ init \ prediction)$$

$$N_{mod}^{exp} = \sum_{bins} \int_{bin\#i} \phi(E) \sigma(E) \epsilon(E) (1 + f_i) dE$$

$$\sigma_{N_{0,mod}^{exp}}^{det.} = N_{0,mod}^{exp} \sigma^{det.}, \quad \sigma^{det.} : \text{common among modules}$$

# Next step

- Proceed this analysis with Fake MC (to understand fitting).
- Consider treatment of some systematic errors in the analysis.
  - Detector systematic.
  - Flux systematic
- Current analysis, the main of neutrino interaction mode is CC int. (~90%). But NC int. remains ~10%.
  - Measured CC int. Cross section in bin#i = Cross section(MC)\_i × (1+f\_i) × Purity\_i (=CC/(CC+NC) in bin#i).
  - Need to consider the uncertainty of Purity\_i (related to neutrino interaction model).

# Back up

# Energy dependence of CC Purity

Purity at module#3 (10d-nominal, NEUT5.0.6)

