

The NASA Space Radiation Laboratory at Brookhaven National Laboratory

K.A. Brown

Collider Accelerator Department, BNL

Outline

The NASA/BNL Space Radiation Program

Sampling of experiments

The NSRL Facility

Operations

Beam Characteristics

Uniform Beams

Dosimetry

Large Dynamic Range Camera Imaging System

Solar Particle Simulator

Summary

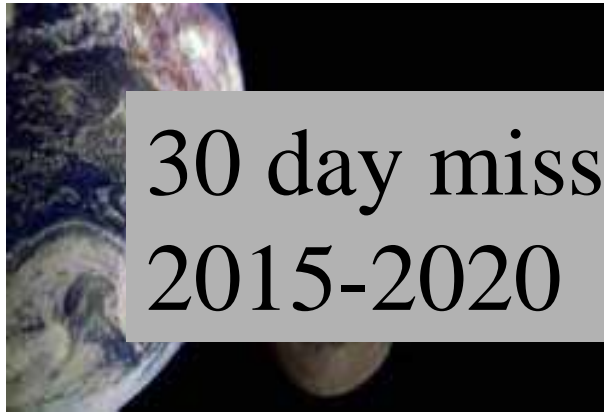
The NASA/BNL Space Radiation Program

“Because astronauts are spending more time in space, the National Aeronautics and Space Administration (NASA) is working with Brookhaven National Laboratory and others here on Earth to learn about the possible risks to human beings exposed to space radiation. To study the radiobiological effects using beams that simulate the cosmic rays found in space, a new, \$34-million NASA Space Radiation Laboratory (NSRL) has been established at Brookhaven Lab.”

http://www.bnl.gov/medical/NASA/NSRL_description.asp

Human Environments in Space

Present



Future



Radiation Doses on Earth and in Space

| | |
|---------------------------|---|
| 1y in Houston | 100 mrem |
| 1y in Denver | 200 mrem |
| 1y in Kerala, India | 1,300 mrem |
| Apollo 14 | 1,100 mrem (9-day to the Moon) |
| Skylab 4 | 18,000 mrem (87-day in orbit) |
| Shuttle mission 41-C | 5,600 mrem (18-day in orbit) |
| Mission to Mars | 130,000 mrem (30 month) |
| at solar minimum | 30,000 mrem in 1.5 y on Mars 80,000 mrem in 1 y in space, +20,000 mrem from a solar flare |
| Chest X-ray | 50 mrem |
| PET scan | 1,000 mrem |
| Treatment of brain cancer | 500,000,000 mrem (to normal brain) |

The NASA/BNL Space Radiation Program

NSRL became operational during summer 2003

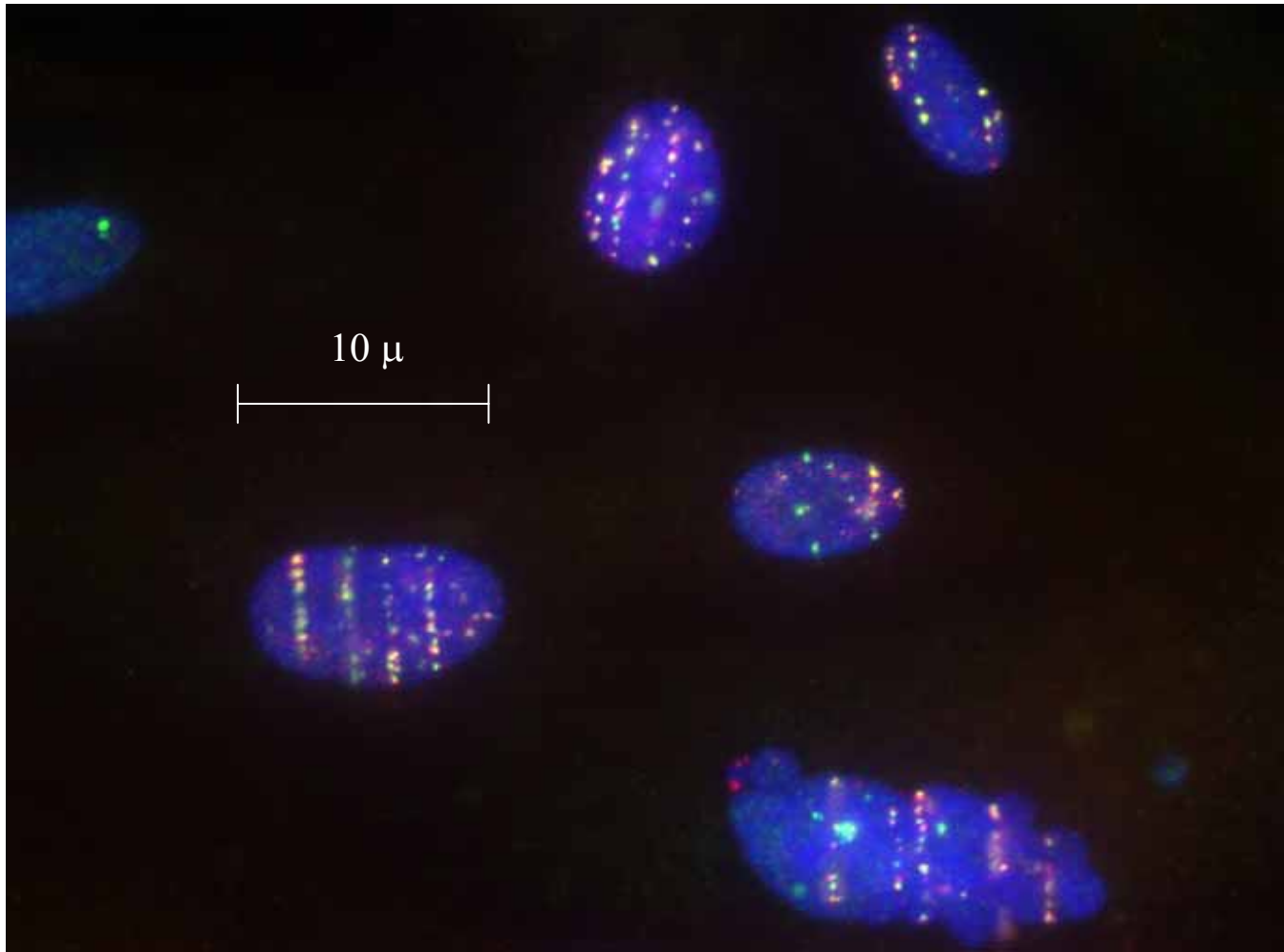
> 100 experimenters from 24 institutions (U.S. and abroad)

Brookhaven researchers and other NASA-sponsored scientists irradiate a variety of biological specimens, tissues, and cells, as well as DNA in solution. Other experimenters use industrial materials as samples, studying their suitability for space suits and spacecraft shielding.

Sampling of experiments

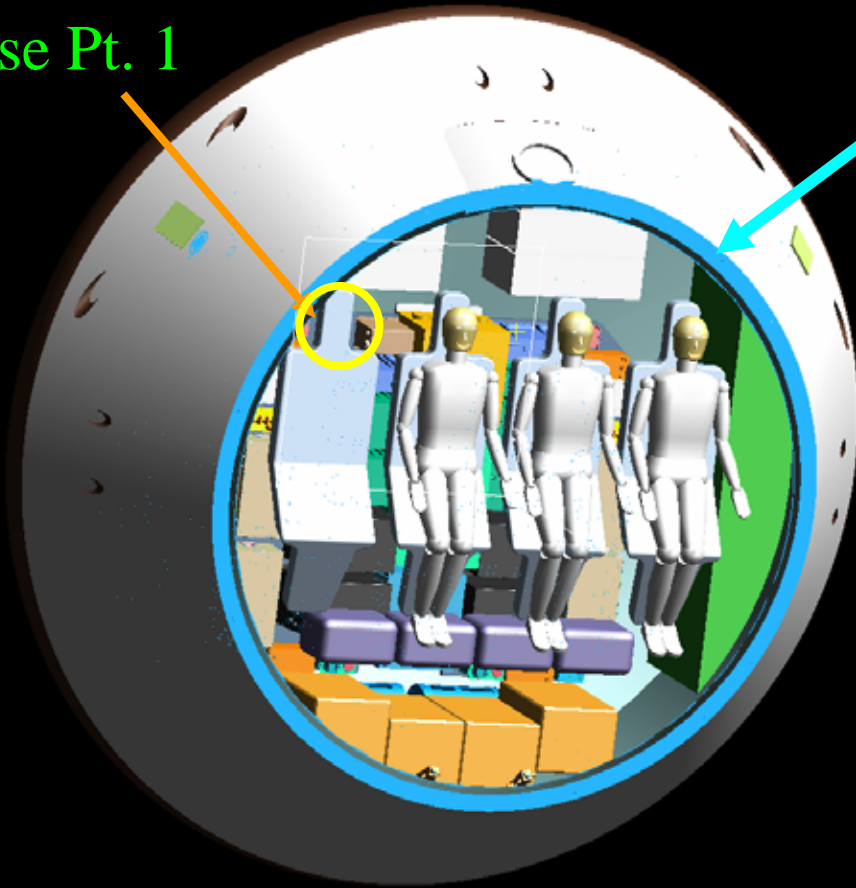
- Effect of Deep Space radiation on Human Hematopoietic Stem Cells
- Risk Assessment and chemoprevention of HZE-induced CNS damage
- Heavy Ion Induced Chromosome Damage and Biomedical Countermeasures
- DNA damage clusters in low level radiation responses of human cells.
- Complex Space Radiation-induced DNA damage Clusters in Human Cell Transformation: Mechanisms, relationships and Mitigation.
- Induction of Bystander Effects by High LET Radiation in Cells
- Gene Expression in the Nematode *C. elegans* following Irradiation with Charged Particles
- Heavy Ion Particle Impact on Simulated Martian Regolith
- MSL/RAD Technology Demonstration Model Characterization
- Spacecraft shielding and components experiments
- Ion fragmentation experiments

1 GeV Fe tracks in cells.

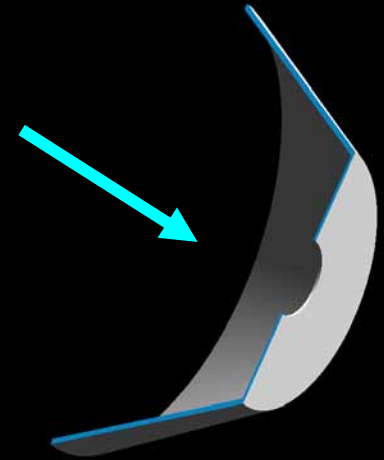


Local Polyethylene Shielding Study

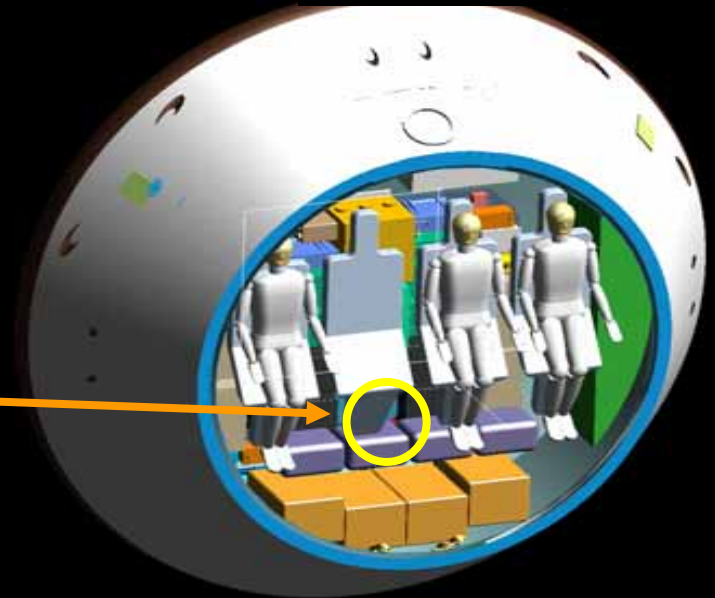
Dose Pt. 1



Polyethylene radiation shield



Dose Pt. 2



Outline

The NASA/BNL Space Radiation Program

Sampling of experiments

The NSRL Facility

Operations

Beam Characteristics

Uniform Beams

Dosimetry

Large Dynamic Range Camera Imaging System

Solar Particle Simulator

Summary

Heavy Ions

Charged

Ionizing Radiation

Outer-Space is full of them

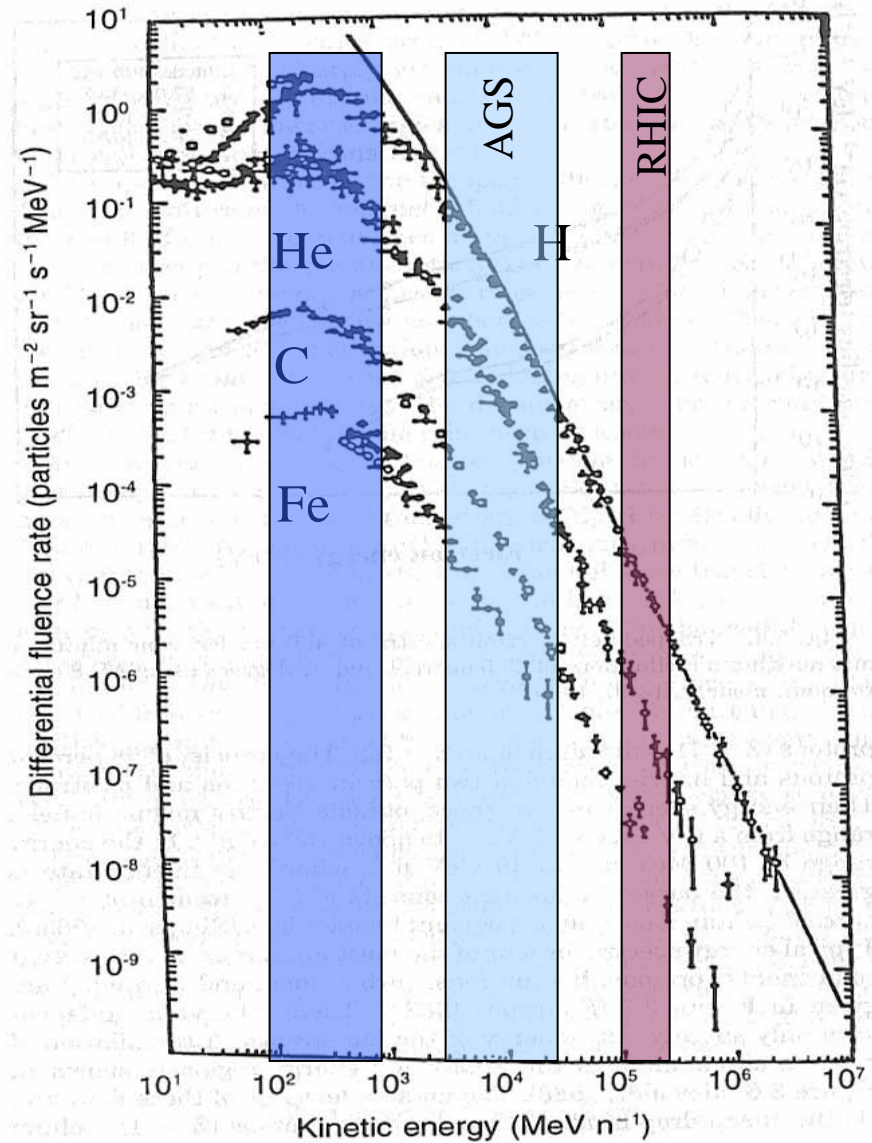
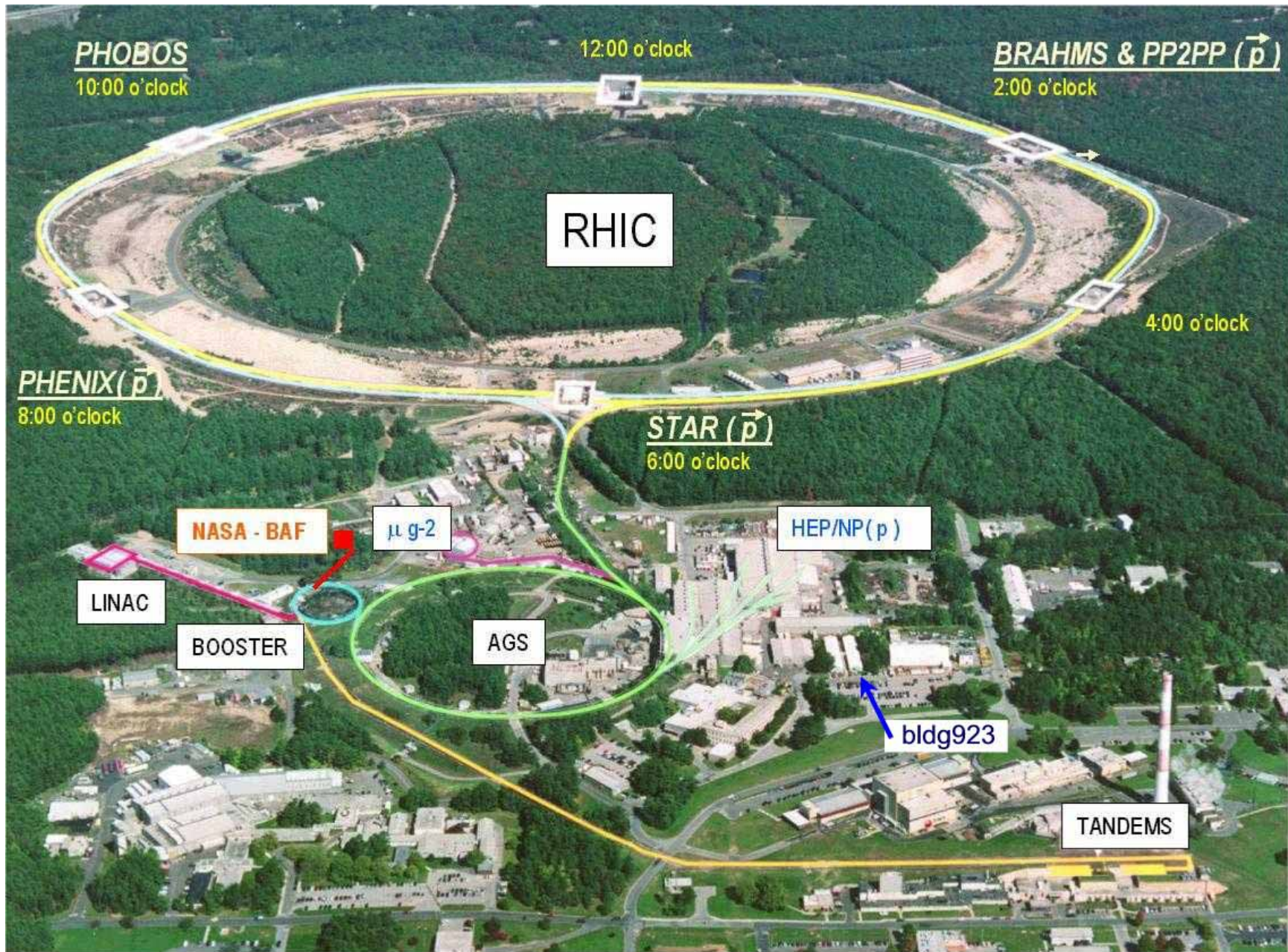
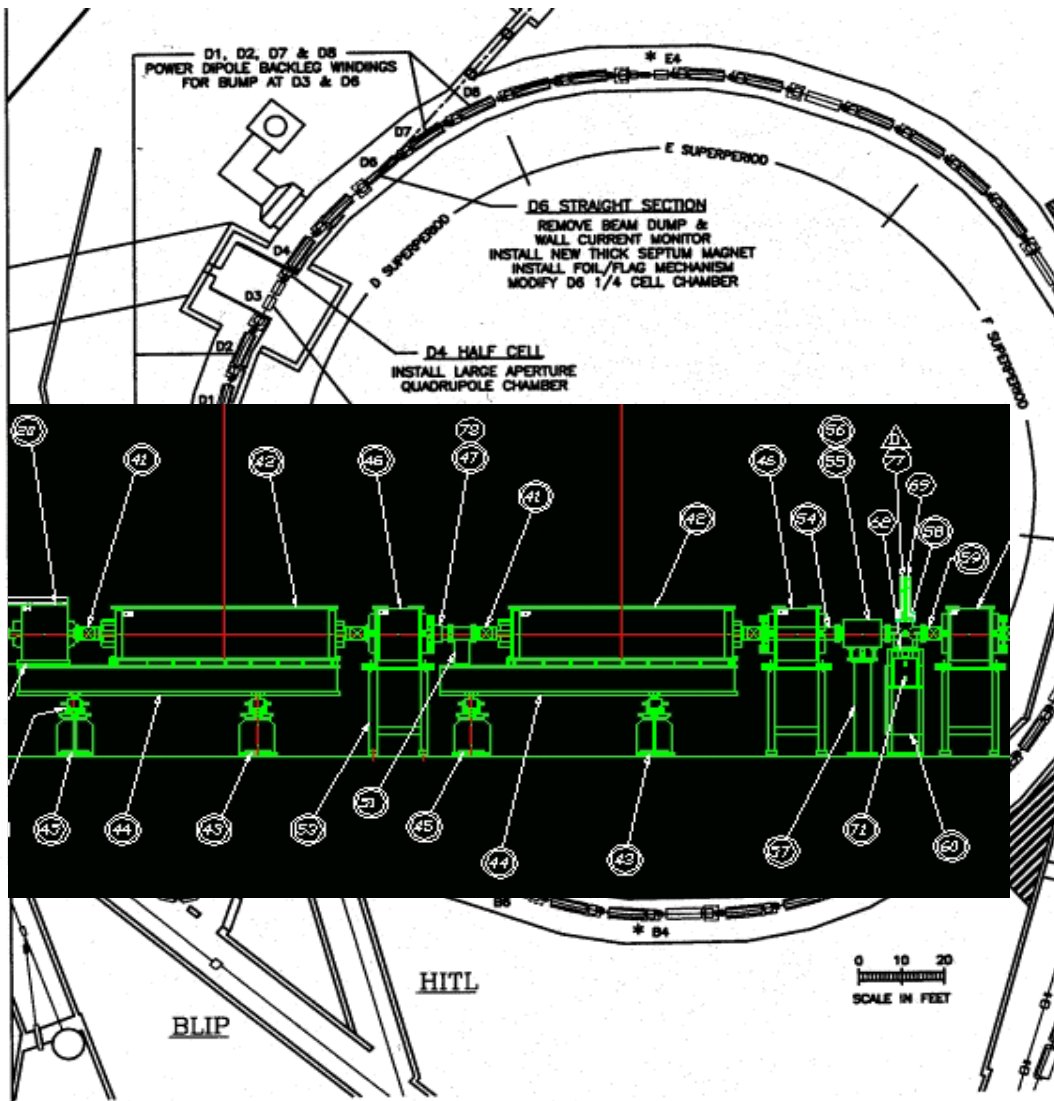


Fig. 3.5. Typical energy spectra for protons, helium ions, carbon ions, and iron ions from “top to bottom,” respectively, at solar minimum. The solid line is the local interstellar spectrum (Simpson, 1983a).

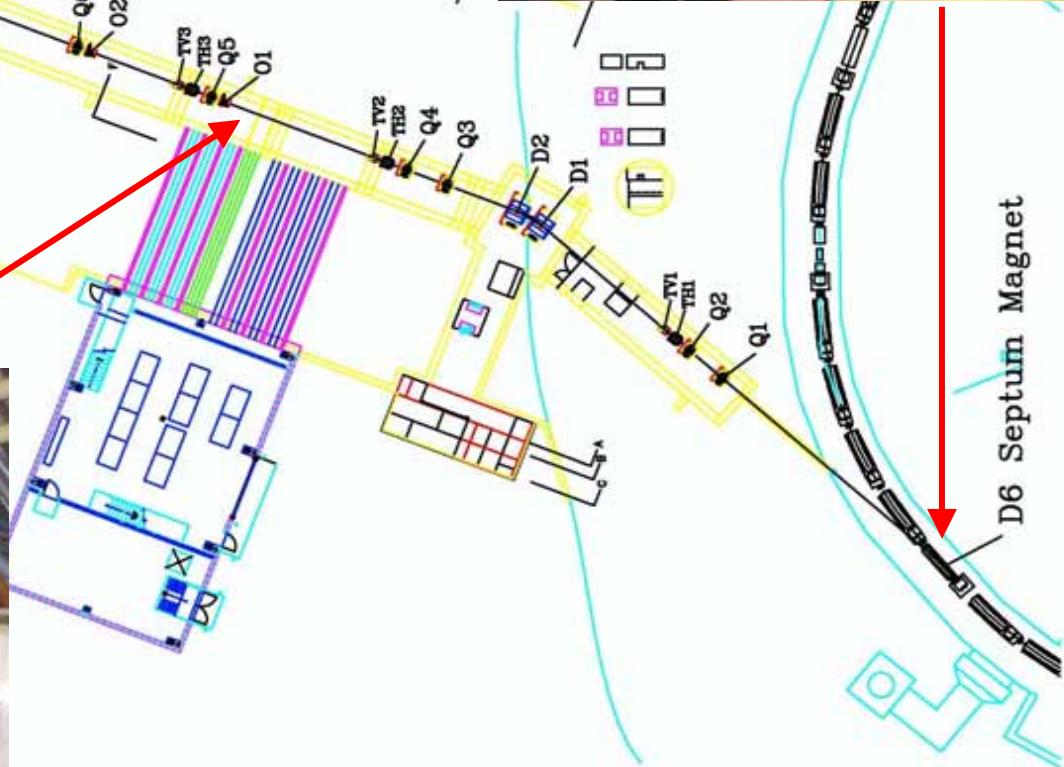
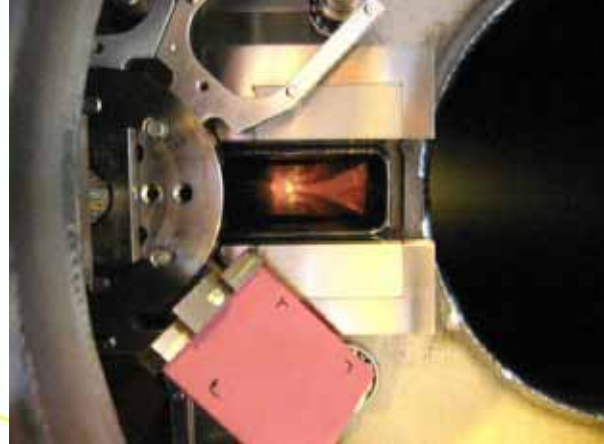
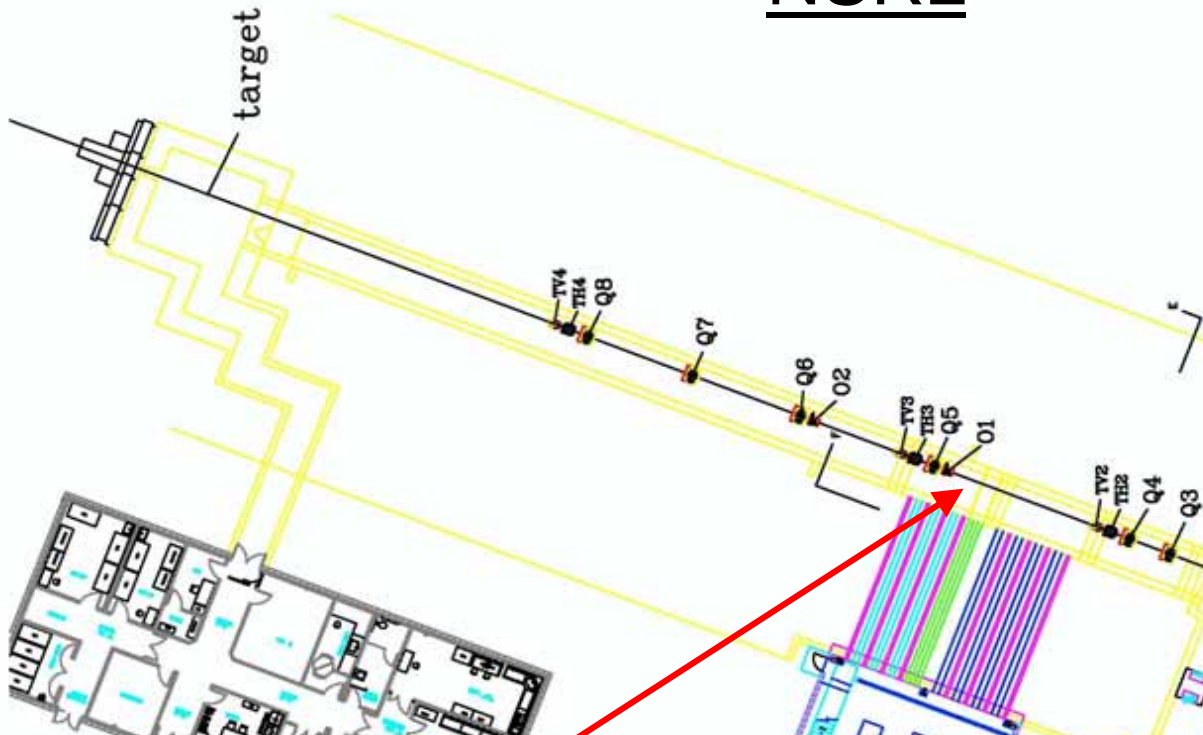


Booster Parameters



| Parameter | Value |
|--------------------|---|
| Circumference | 201.78 (1/4 AGS) m |
| Ave. Radius | 32.114 m |
| Magnetic Bend R | 13.8656 m |
| Lattice Type | Separated Function, FODO |
| No. Superperiods | 6 |
| No. of Cells | 24 |
| Betatron Tunes,X,Y | 4.82, 4.83 |
| Vacuum Chamber | 70 x 152 mm Dipoles 152 mm (circular) Quads |
| Max. Rigidity | 17 Tm |
| Injection Rigidity | 2.2 Tm (200 MeV protons) |
| Acceleration Rate | 8.9 T/s (7.5 Hz) |

NSRL



R-Line: From Booster to NSRL

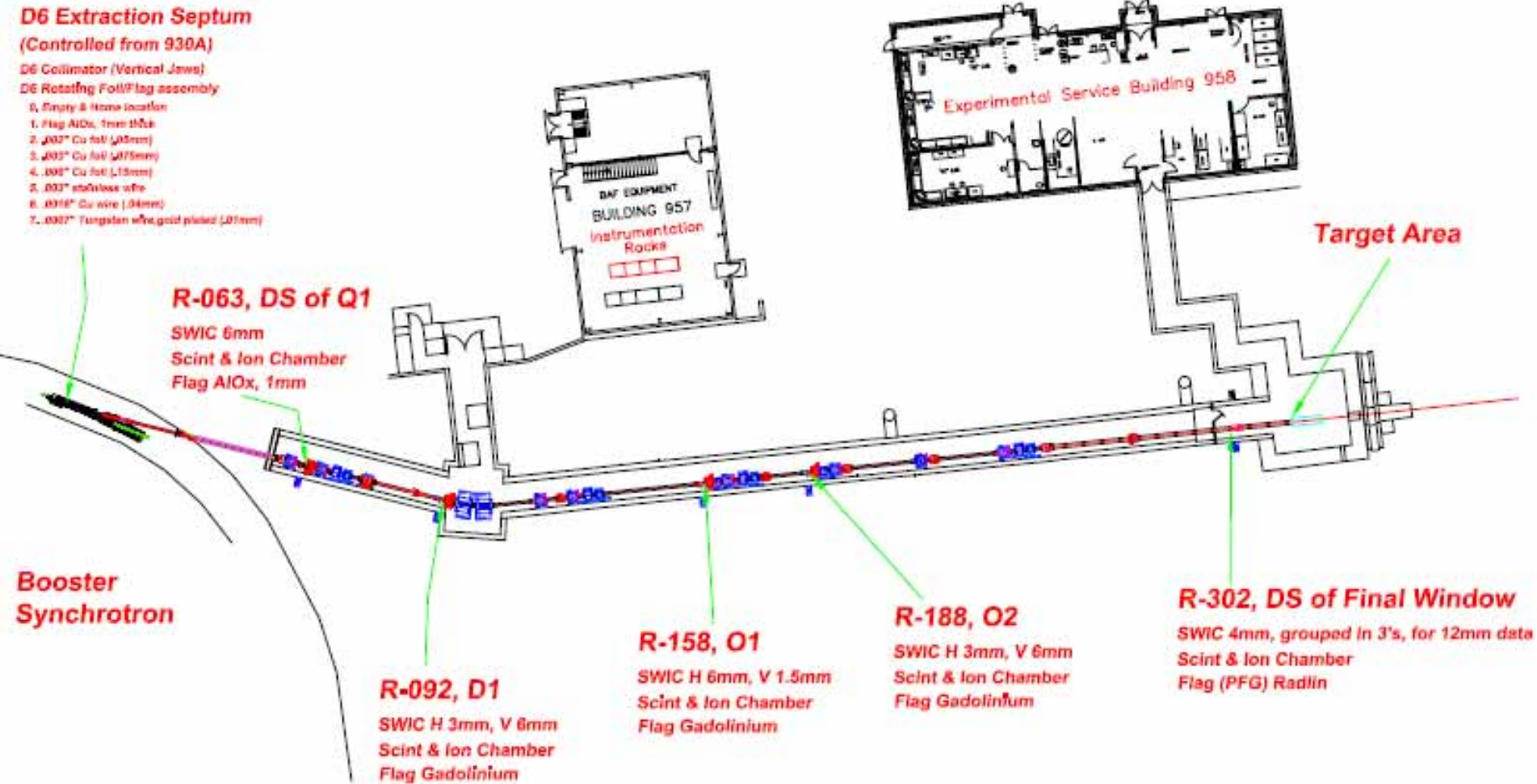


12" beam pipe



8" beam pipe

Instrumentation Layout



NSRL Target Room



Outline

The NASA/BNL Space Radiation Program

Sampling of experiments

The NSRL Facility

Operations

Beam Characteristics

Uniform Beams

Dosimetry

Large Dynamic Range Camera Imaging System

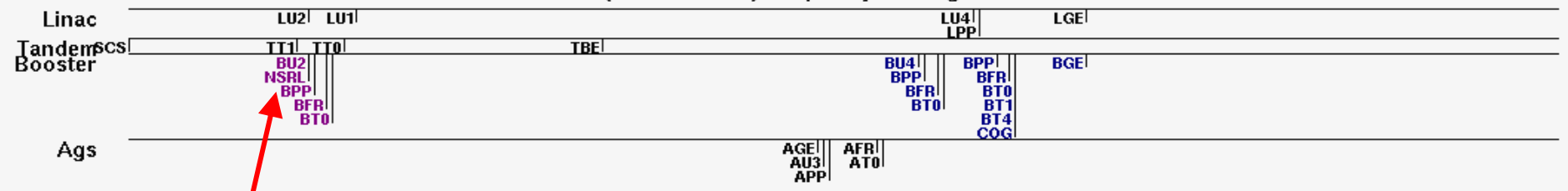
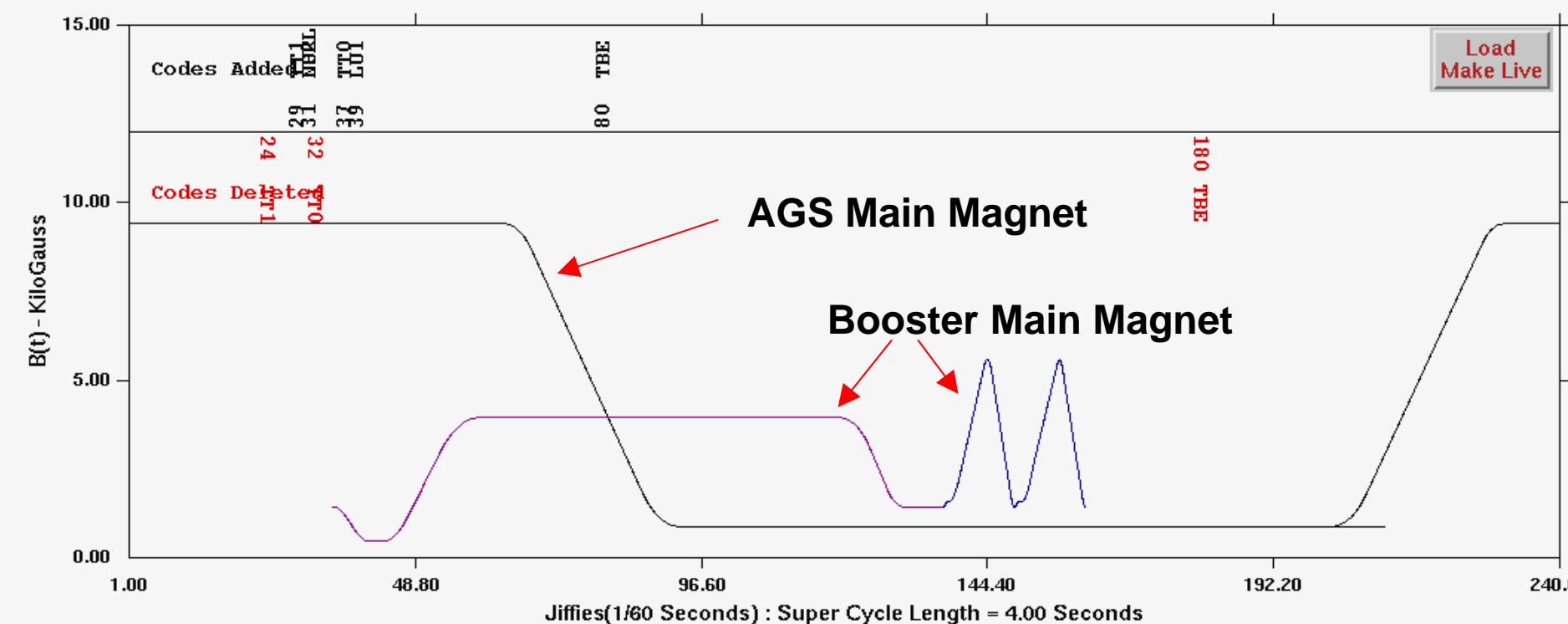
Solar Particle Simulator

Summary

Operations

Beams delivered, NSRL-5 to NSRL-7

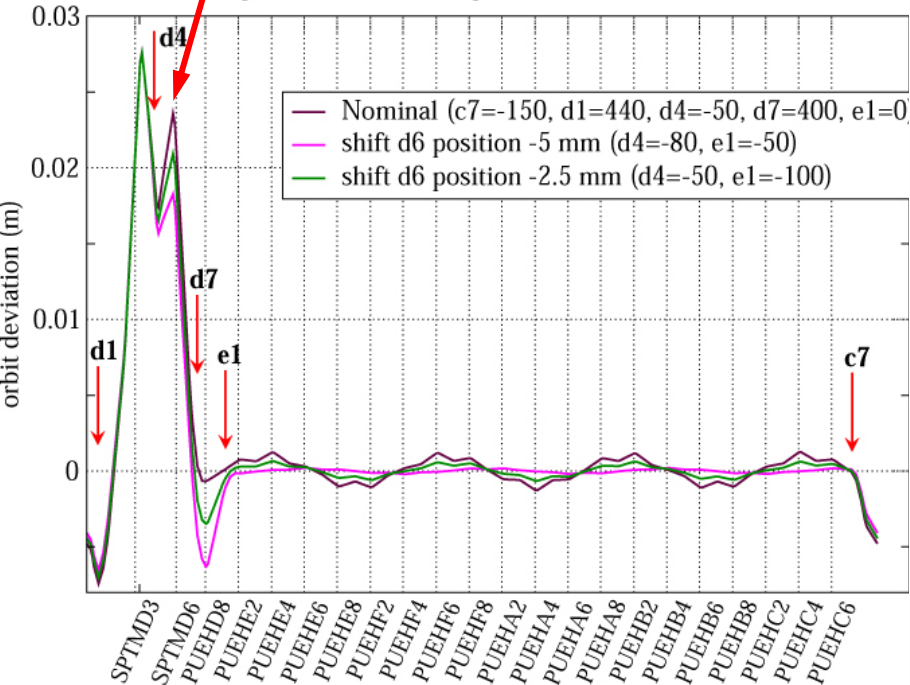
| Ion | Energy (MeV/n) | Intensity |
|-----|----------------|----------------------|
| p | 1000 | 3.4×10^{10} |
| C | 290 | 1.2×10^{10} |
| O | 1000, 600 | 4.0×10^9 |
| Si | 600, 300 | 3.0×10^9 |
| Cl | 500 | 2.0×10^9 |
| Ti | 1100 | 8.0×10^8 |
| Fe | 1000, 600, 300 | 2.0×10^9 |



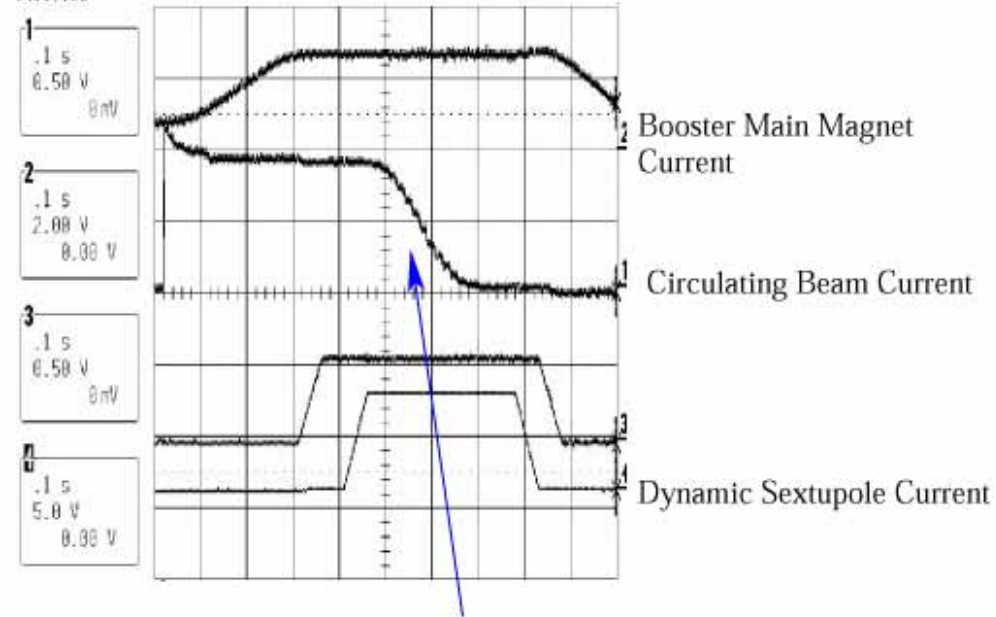
NSRL Event

Extraction Point

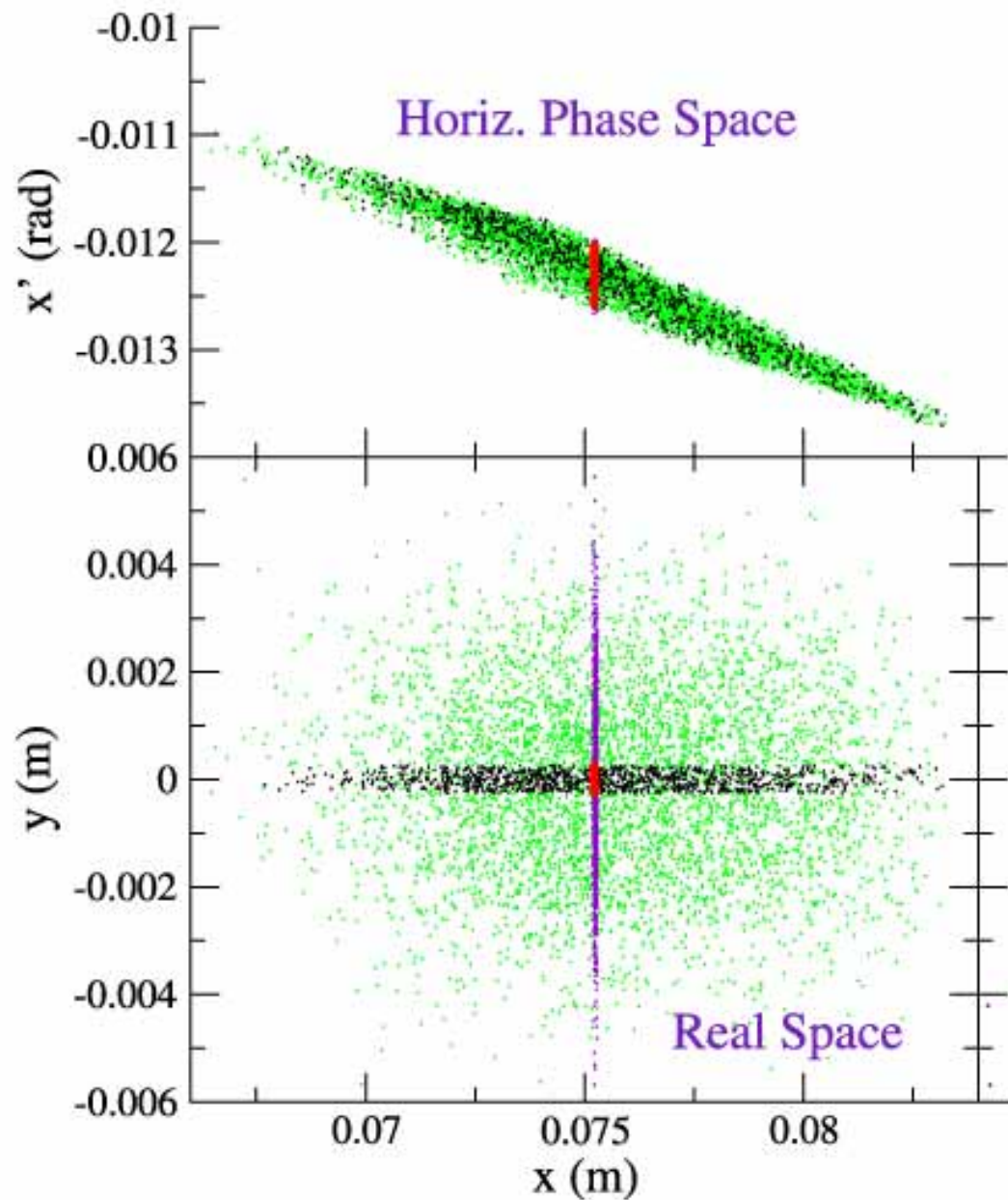
Equilibrium Orbit bump for NSRL slow extraction



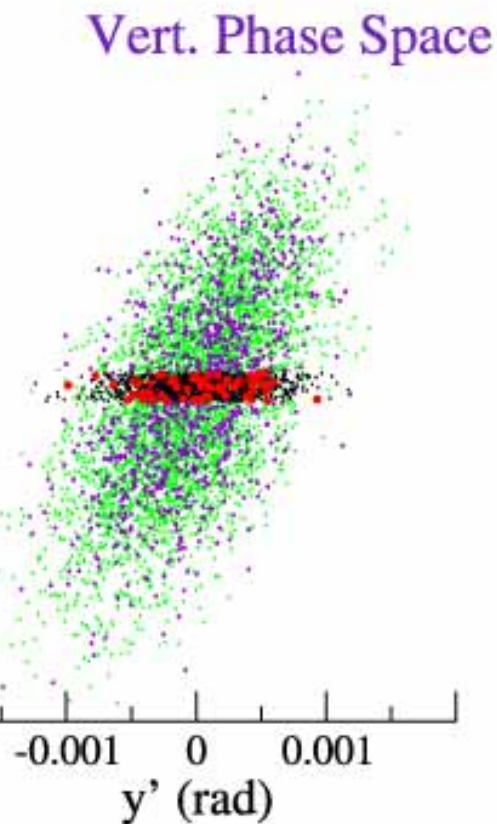
27-Mar-03
14:17:38



Beam Excited by Resonance and Extracting.

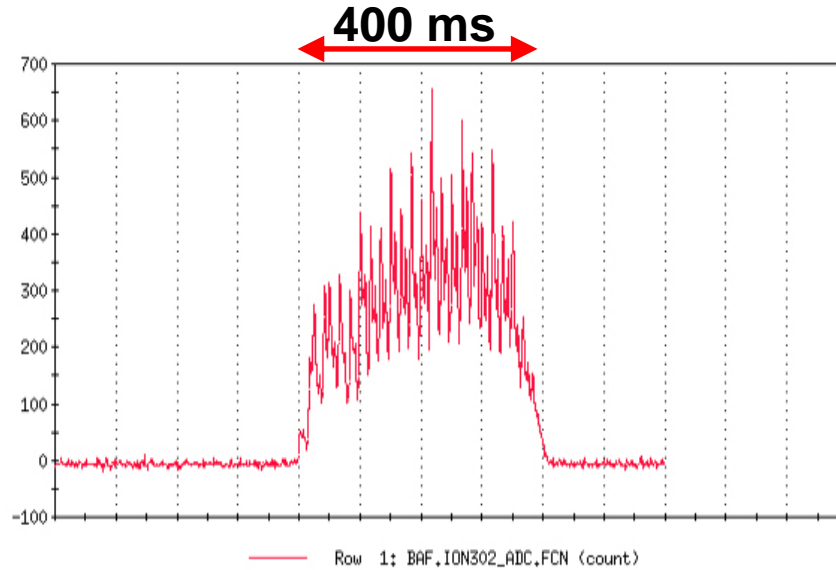
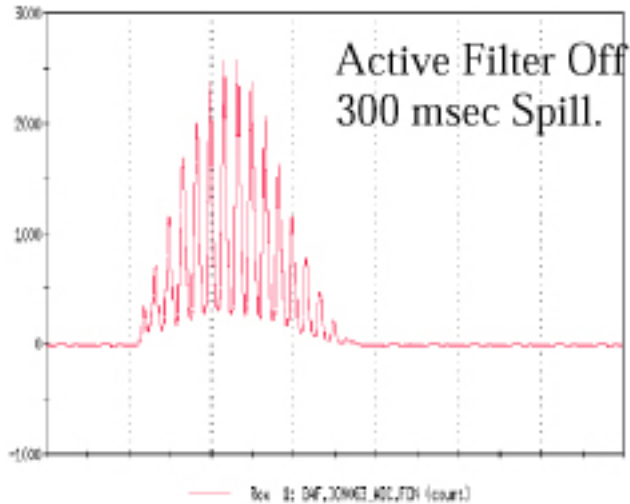


- (green) D6: 0.01 cm full aperture foil, 5k particles
- (violet) D6: wire foil, 100k->1k particles
- (black) D6: collimator 0.5 mm, 10k->1k particles

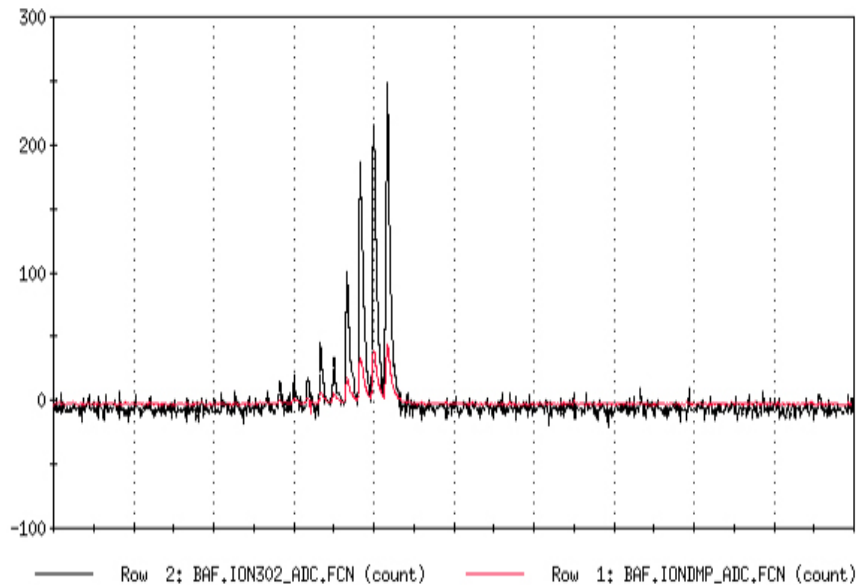
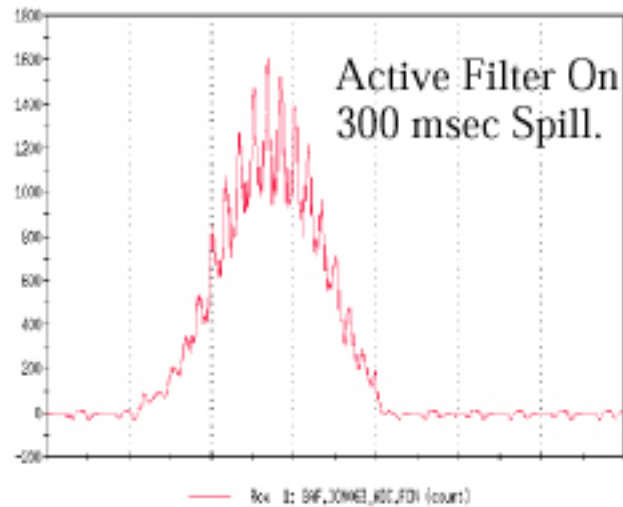


300 MeV/n Carbon

2 GeV Protons

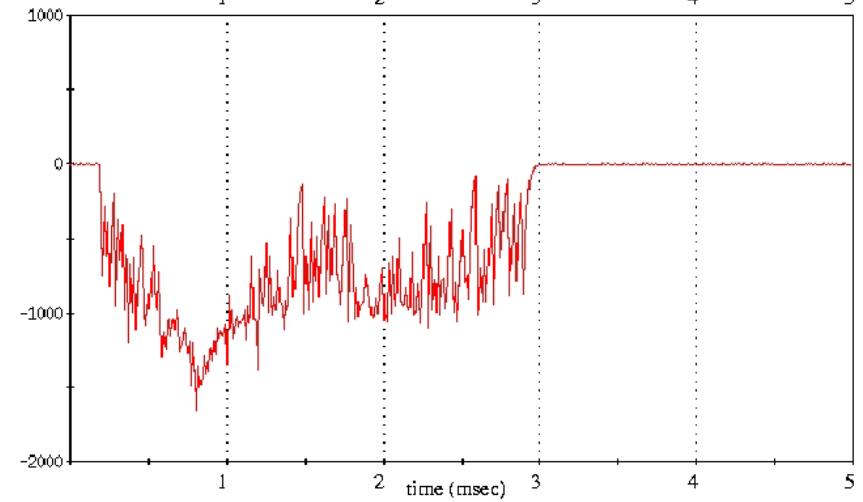
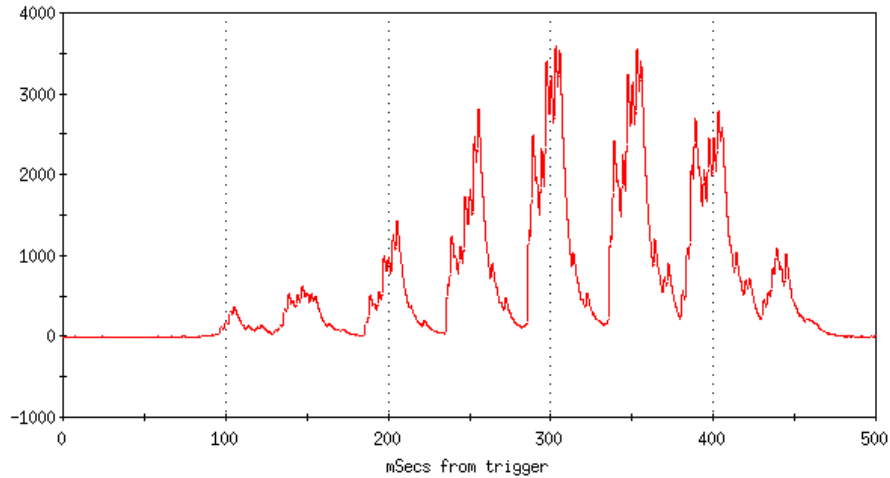
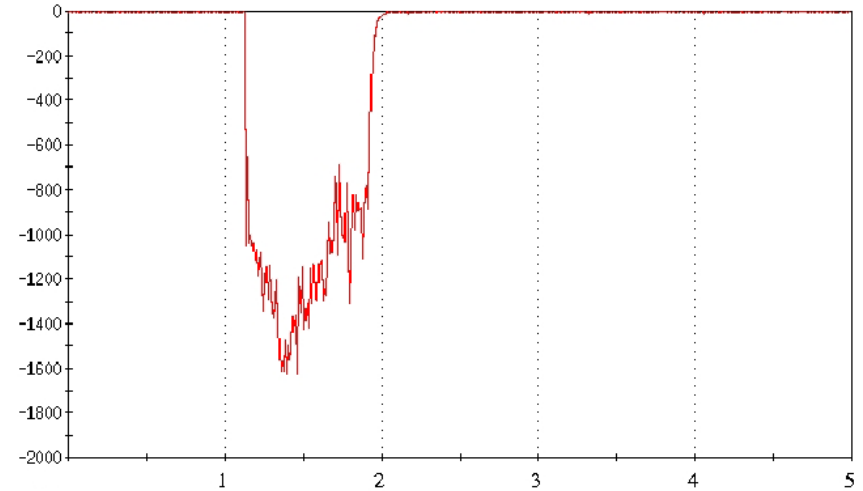
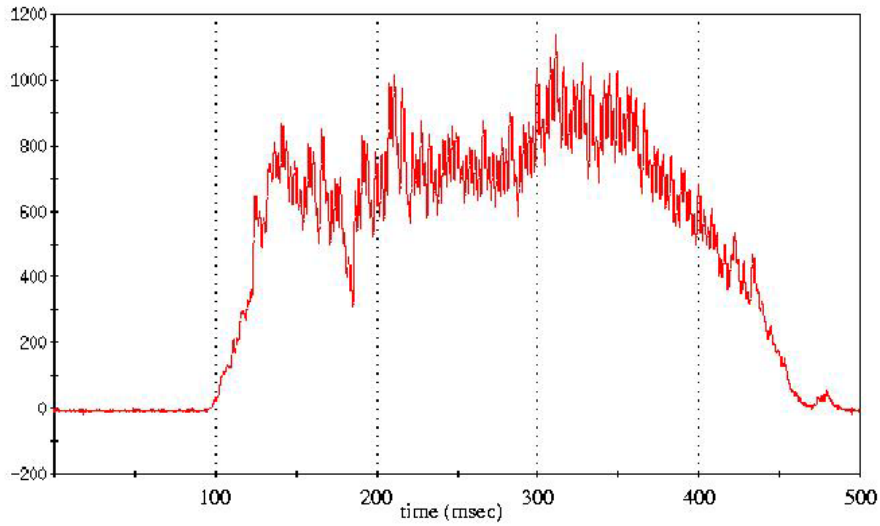


Spill Servo
And Active
Filter On.



Dosimetry
Beam Cutoff

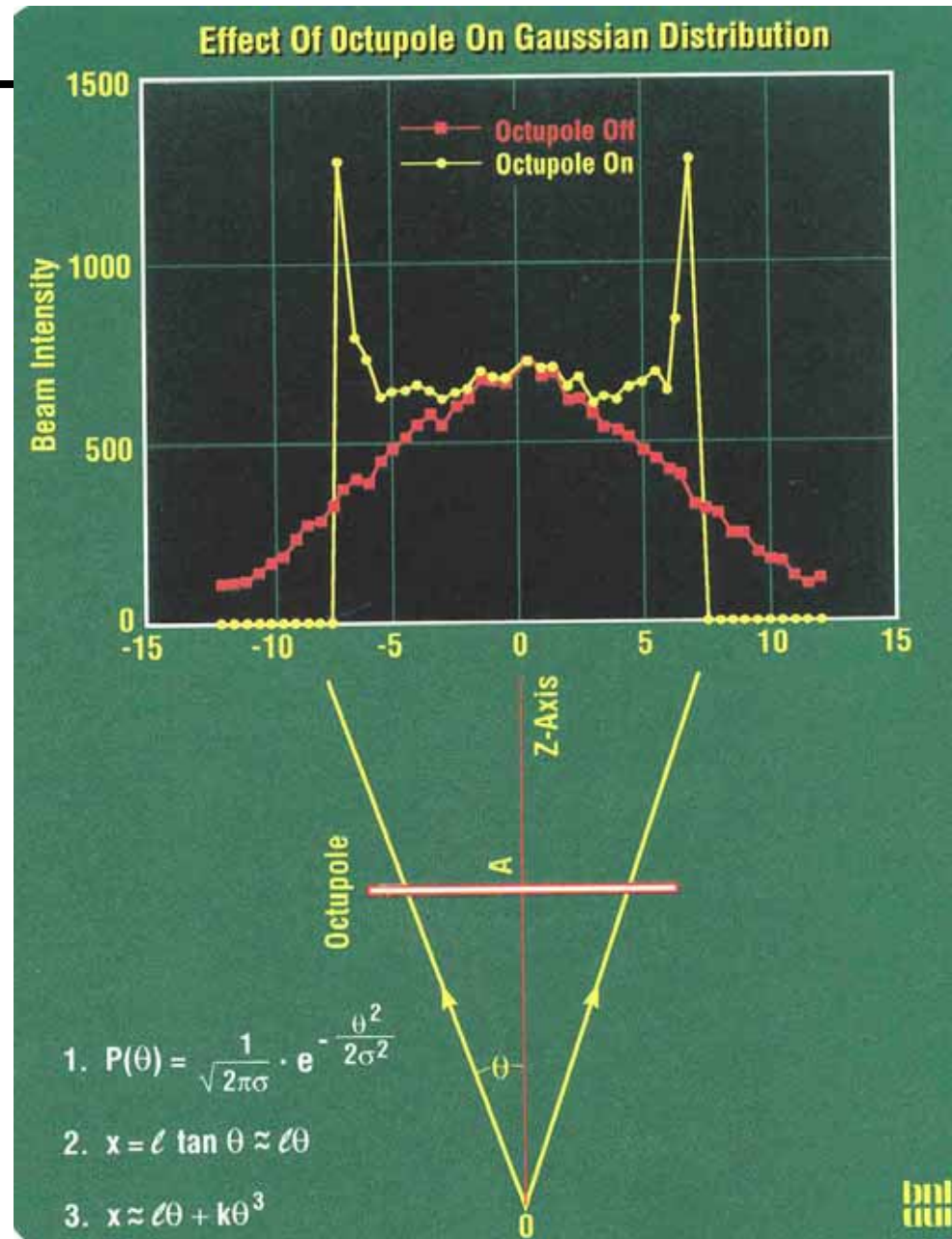
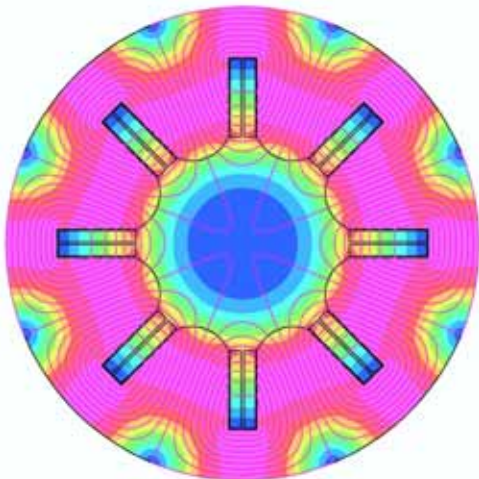
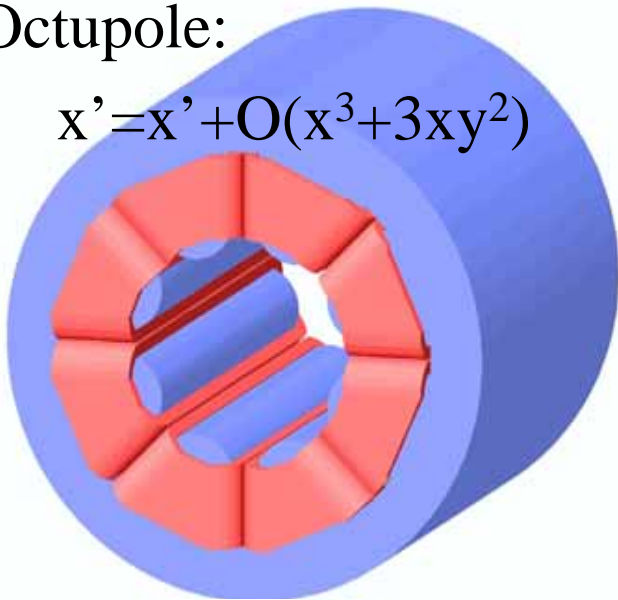
Beam Spills Delivered

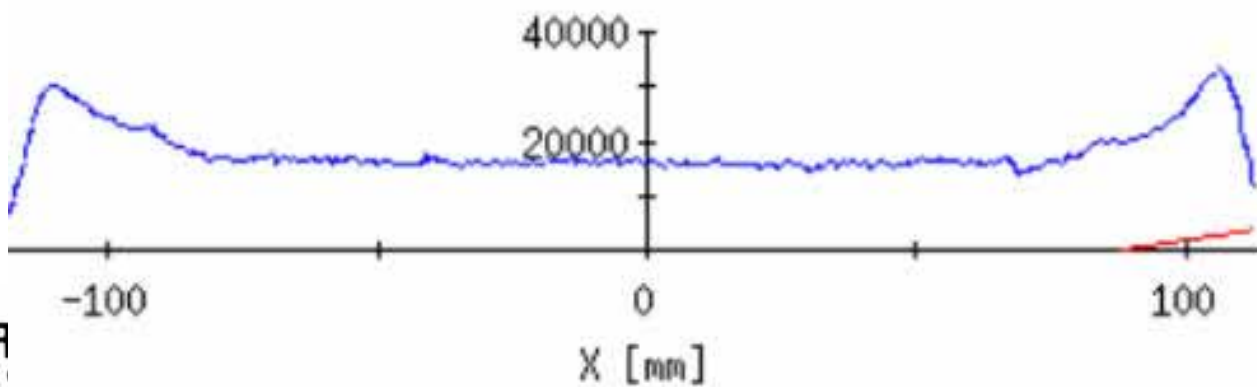
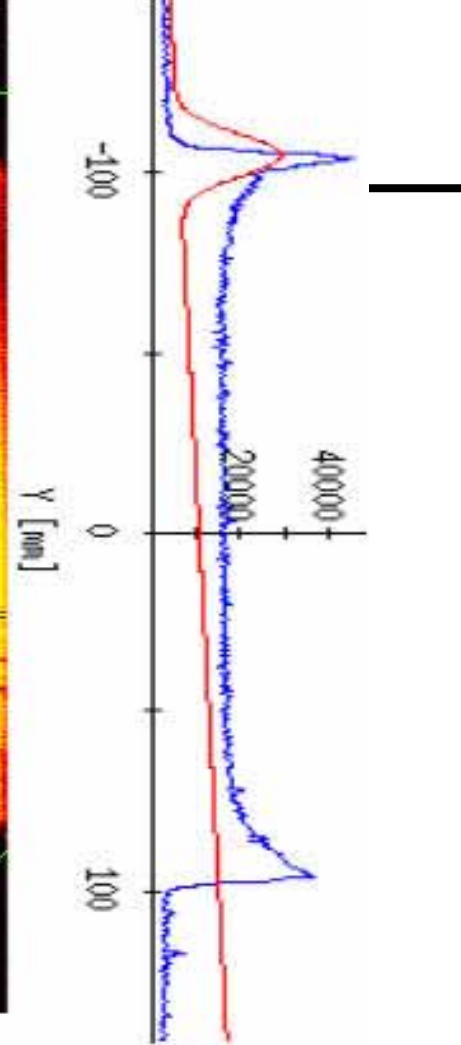
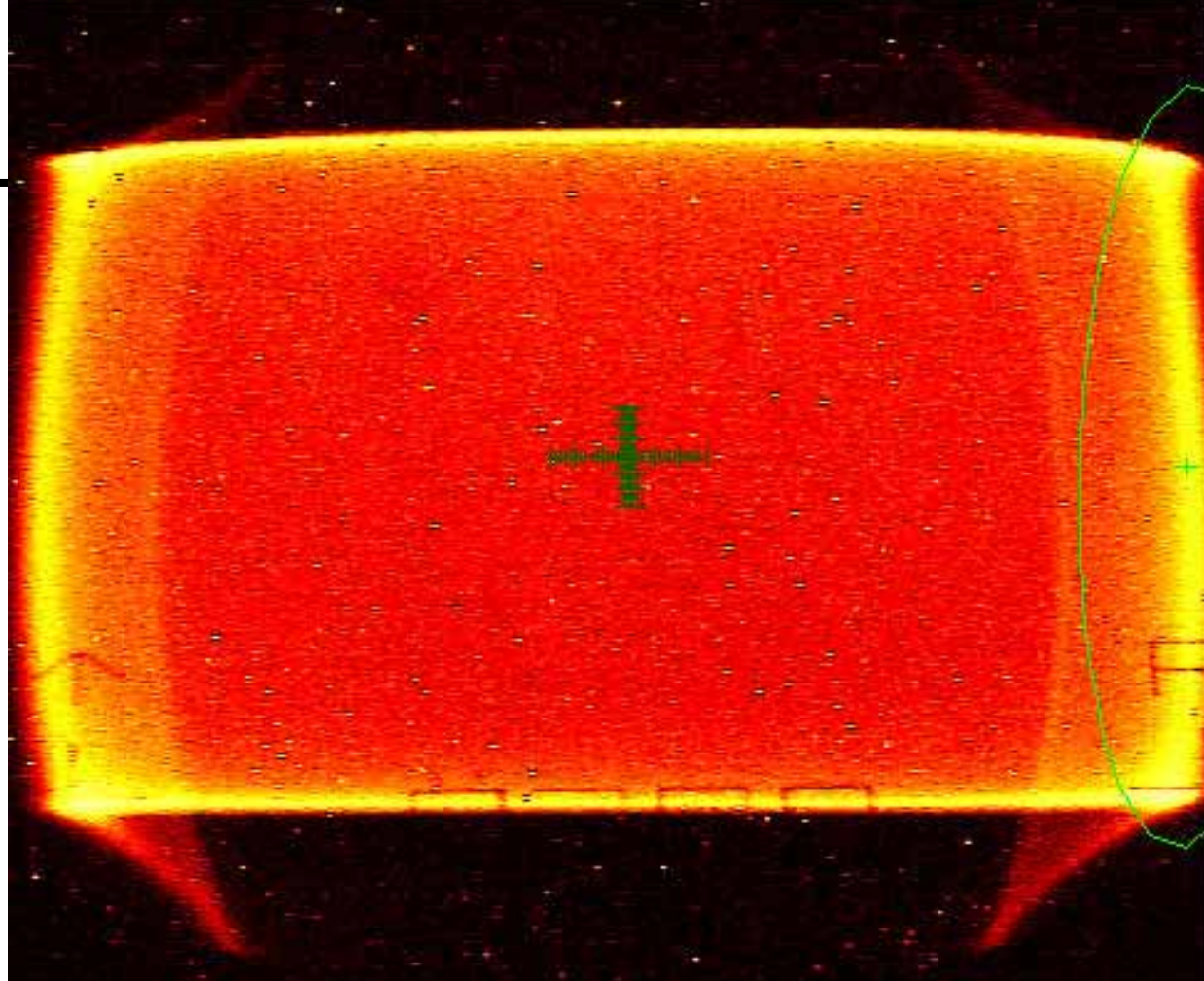


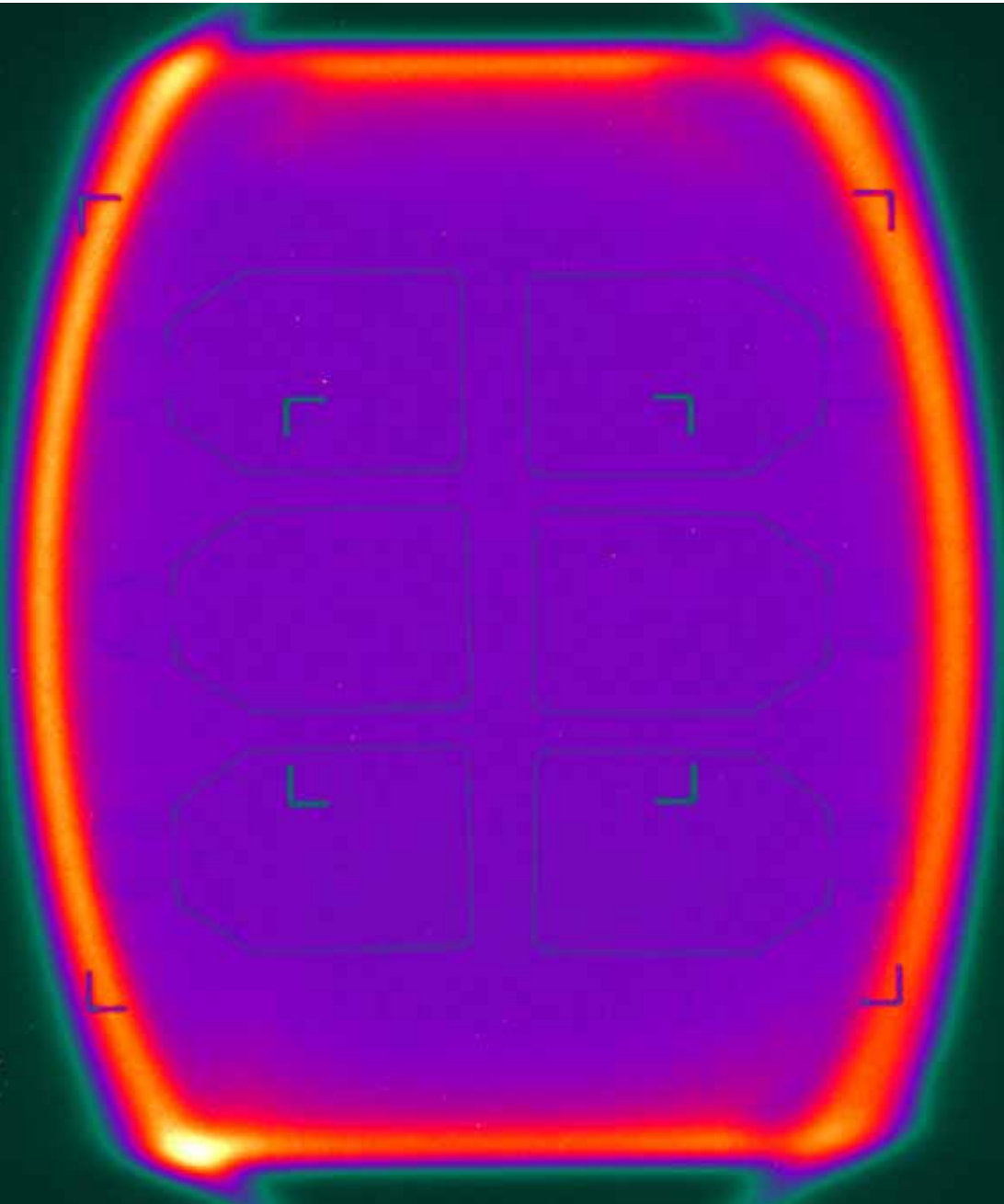
Uniform Beams

Octupole:

$$x' = x' + O(x^3 + 3xy^2)$$







Outline

The NASA/BNL Space Radiation Program

Sampling of experiments

The NSRL Facility

Operations

Beam Characteristics

Uniform Beams

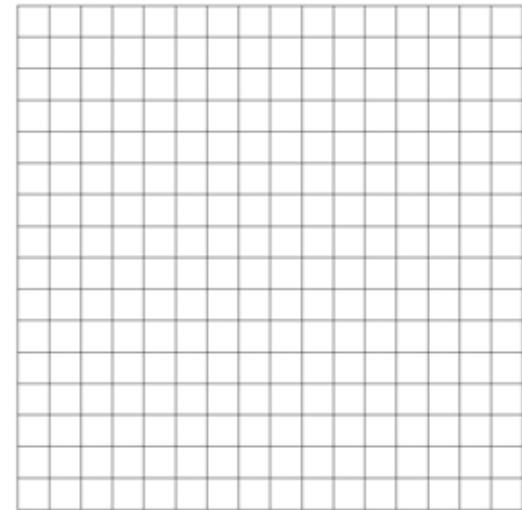
Dosimetry

Large Dynamic Range Camera Imaging System

Solar Particle Simulator

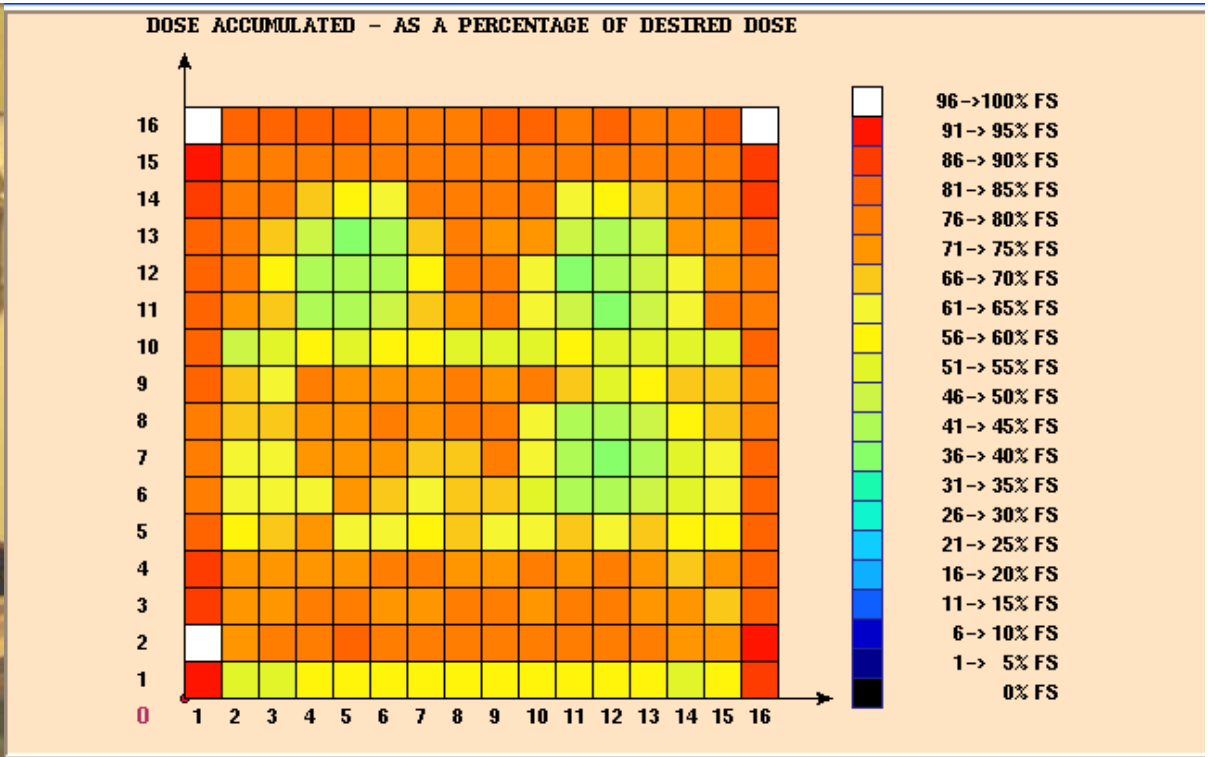
Summary

256-Element Ion-Chamber (Beam Imaging)

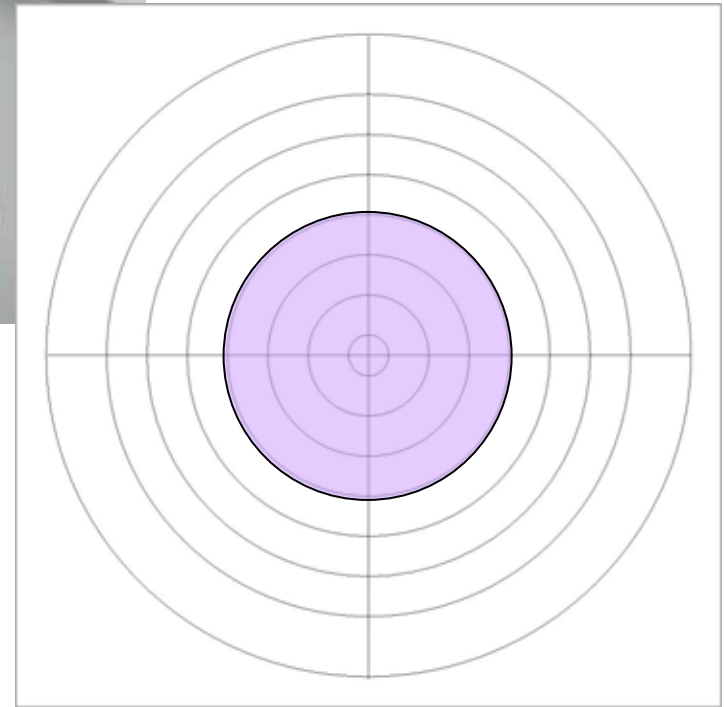


Beam Image on the 256-element IC

During E.Blakely's experiment (NSRL-2).



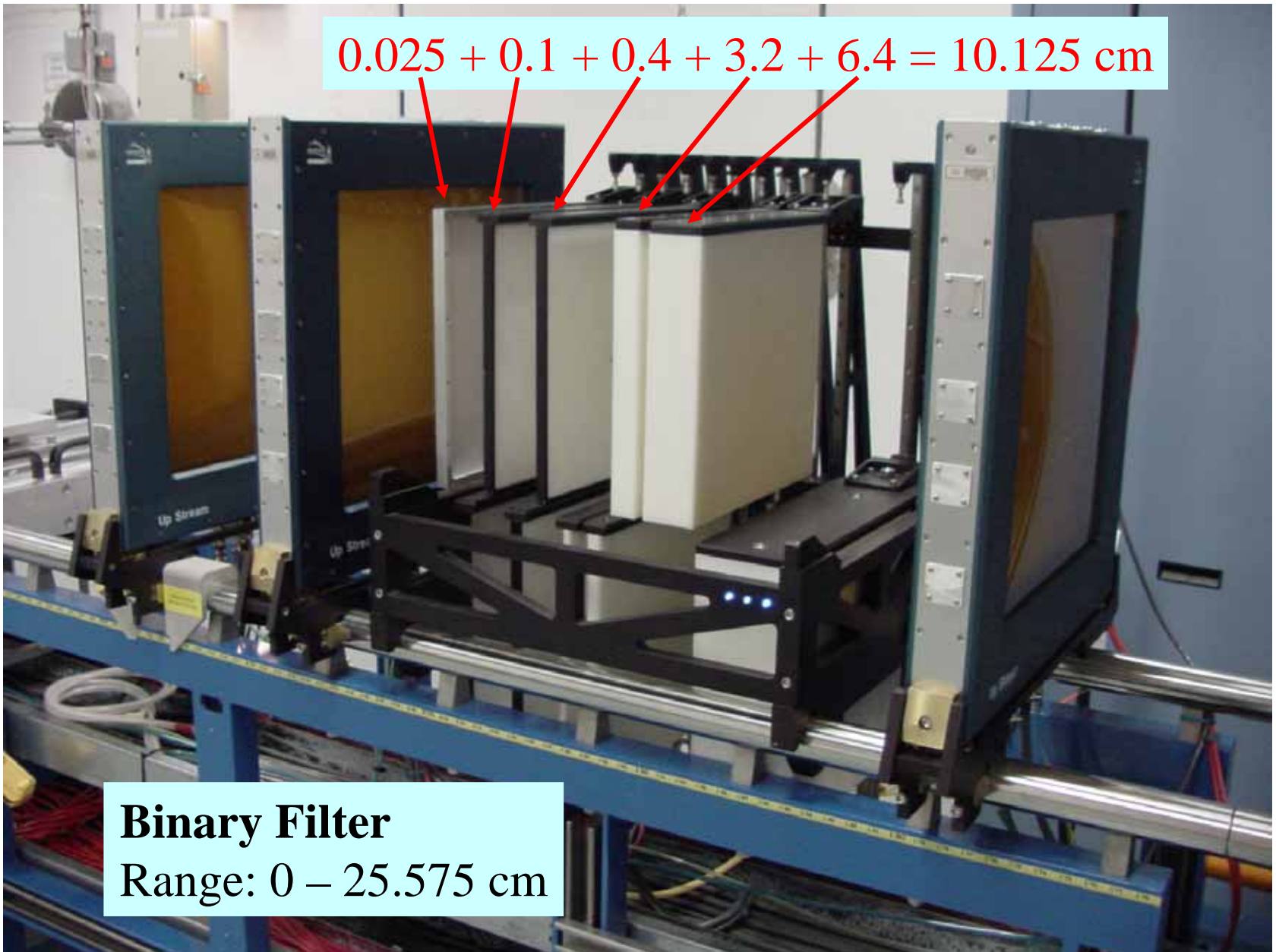
32-Element Ion-Chamber (Dosimetry)



$$0.025 + 0.1 + 0.4 + 3.2 + 6.4 = 10.125 \text{ cm}$$

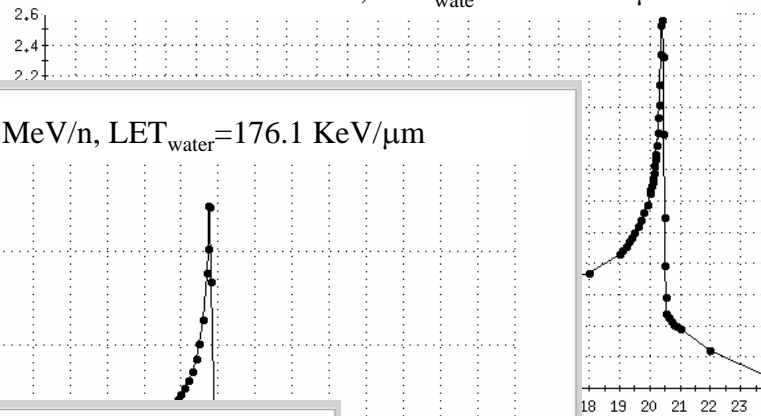
Binary Filter

Range: 0 – 25.575 cm

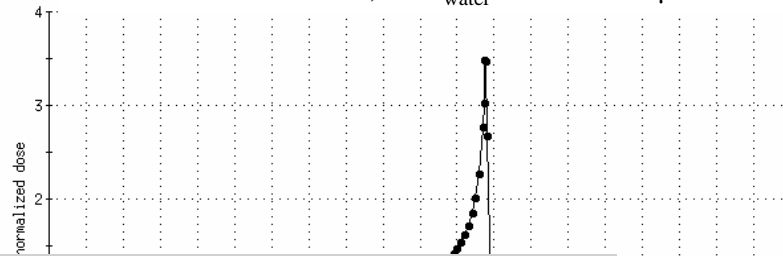


C. KE=294.1 MeV/n, LET_{water}=13.6 KeV/μm

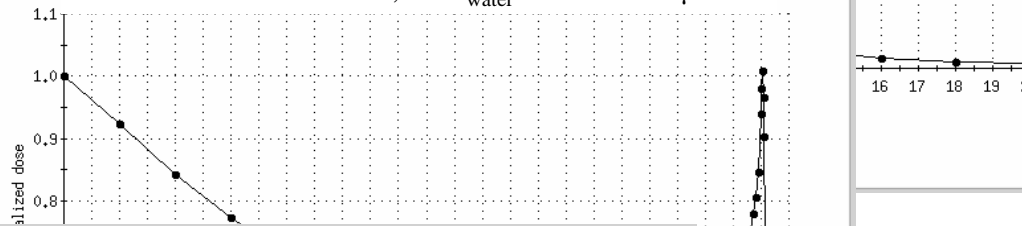
Si. KE=580.6 MeV/n, LET_{water}=53.8 KeV/μm



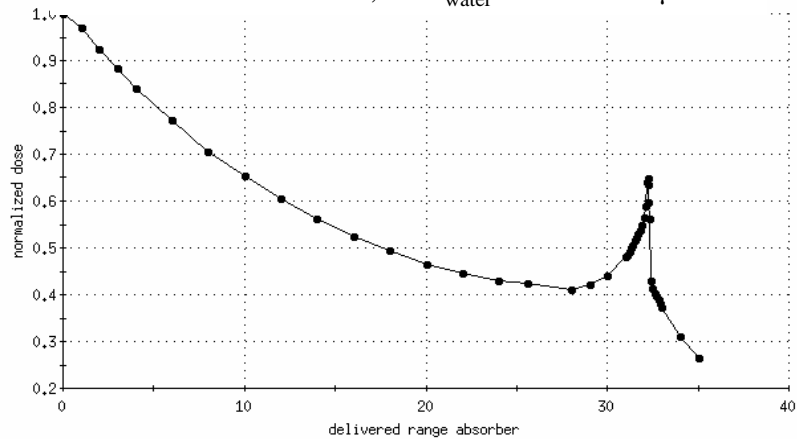
Fe. KE=577 MeV/n, LET_{water}=176.1 KeV/μm



Fe. KE=968.4 MeV/n, LET_{water}=151.4 KeV/μm



Ti. KE=1019 MeV/n, LET_{water}=107.3 KeV/μm



—●— Entire IC

23 24 25 26

16 17 18 19 20

18 19 20 21 22 23 24

14 15 16 17 18 19 20



Adam Rusek



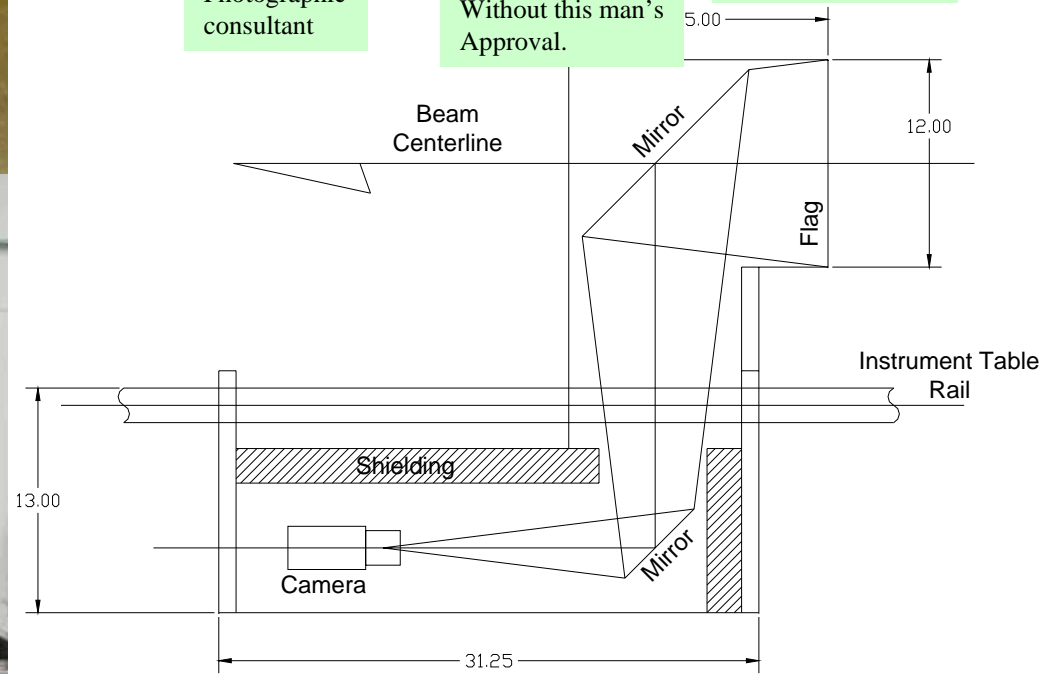
Kevin Brown
Photographic consultant



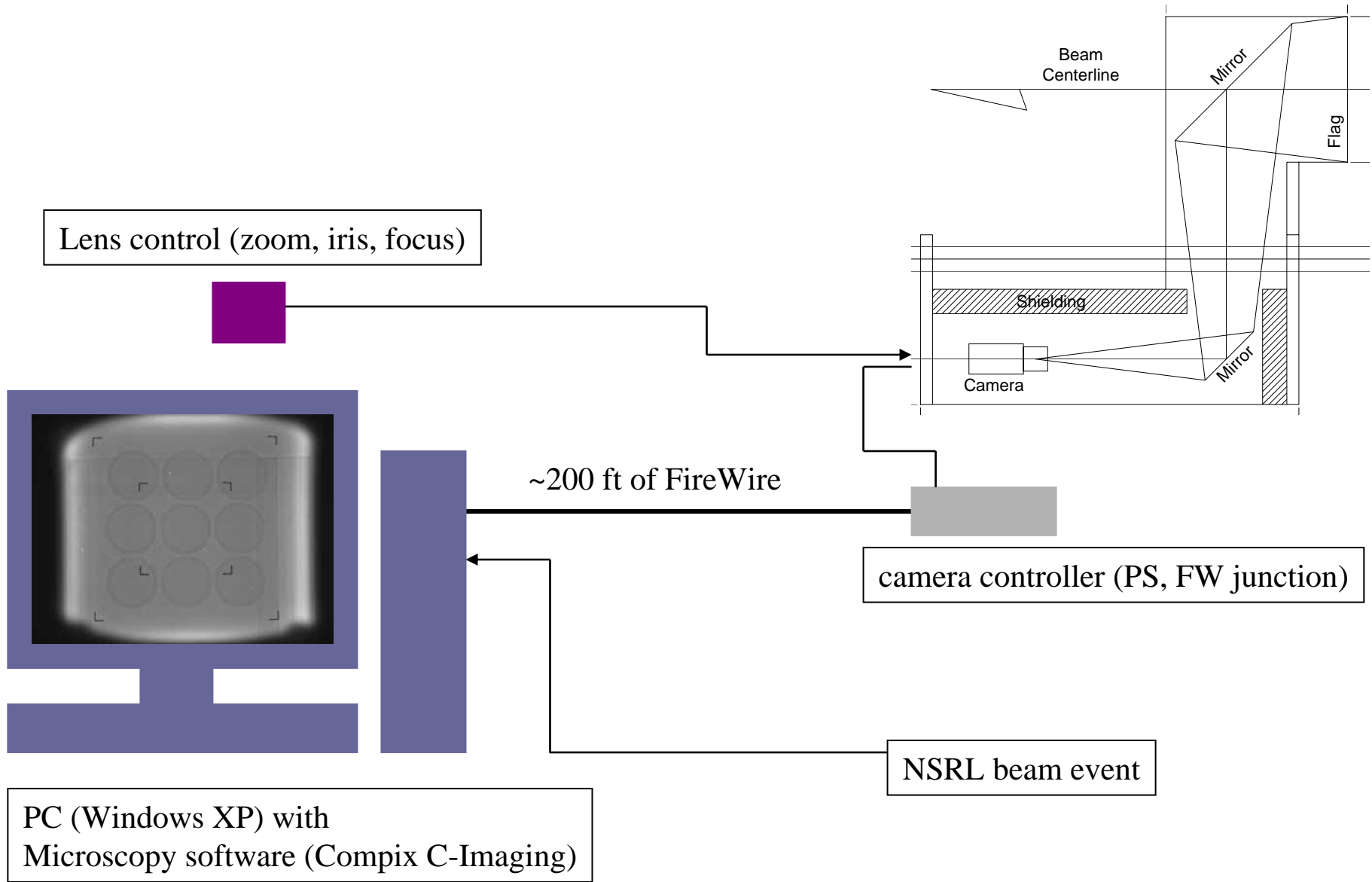
I-Hung Chiang
Nothing happens
Without this man's
Approval.

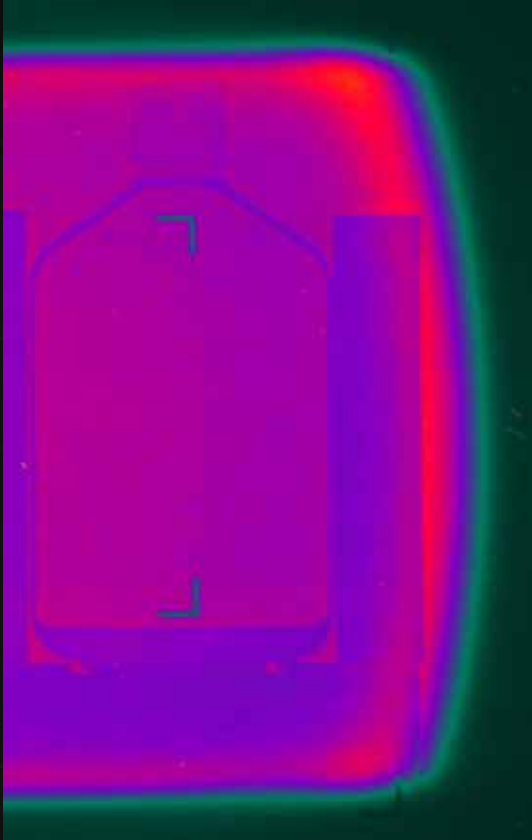
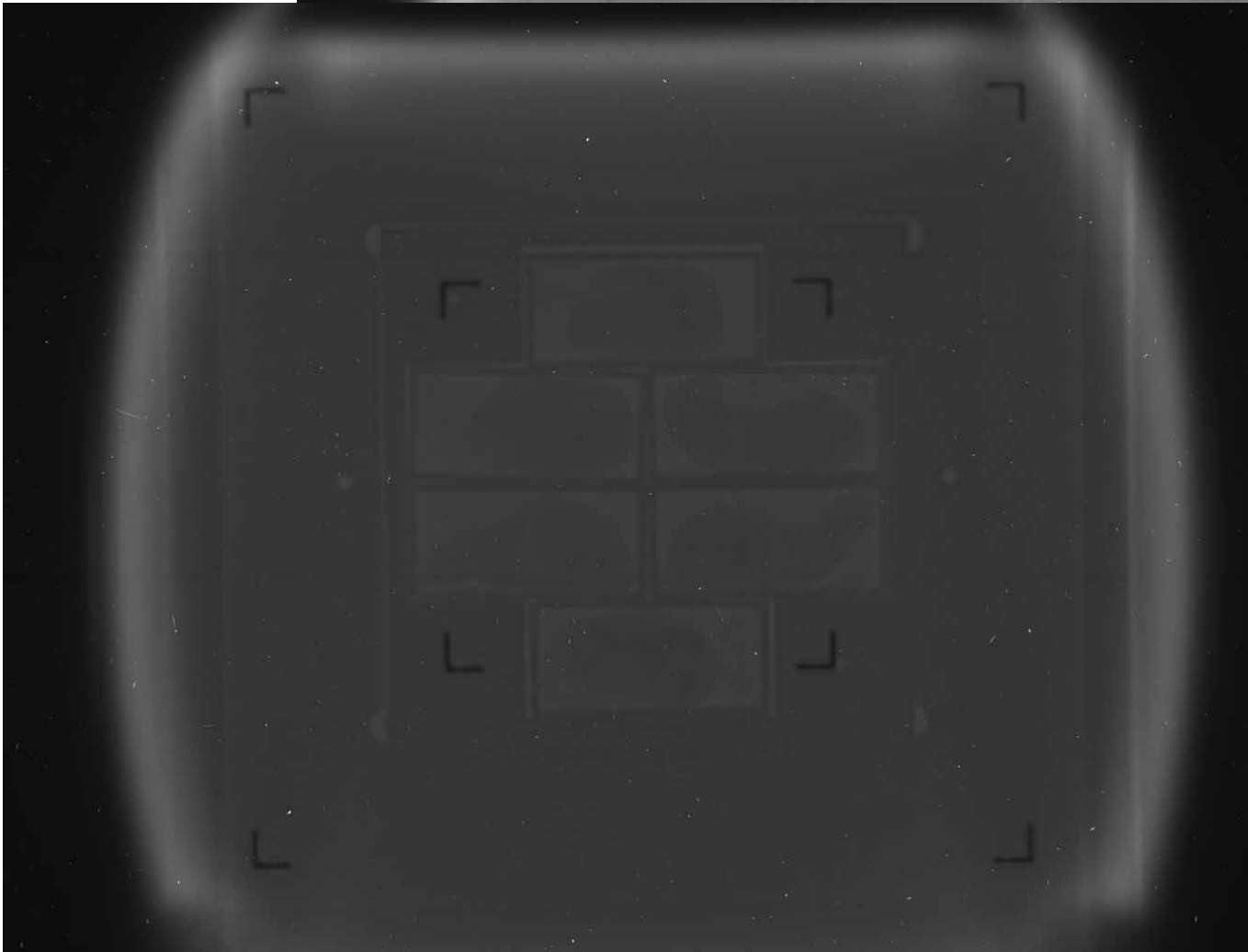
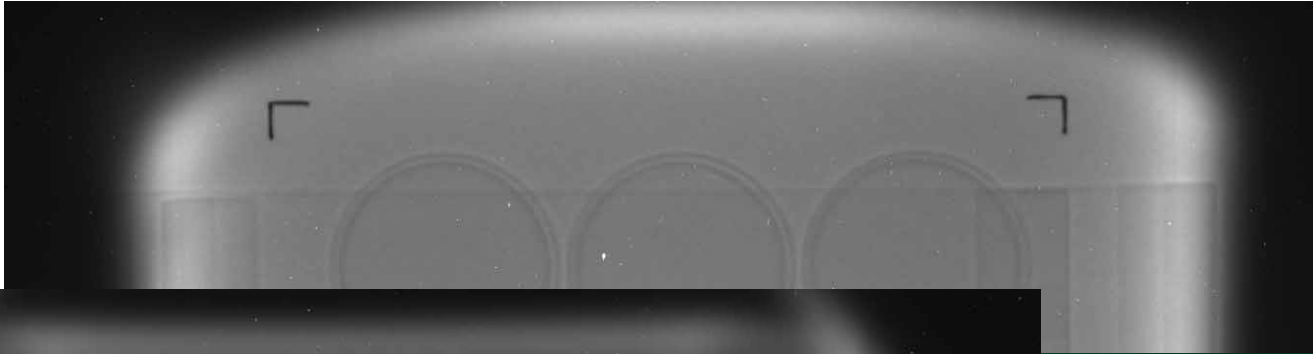


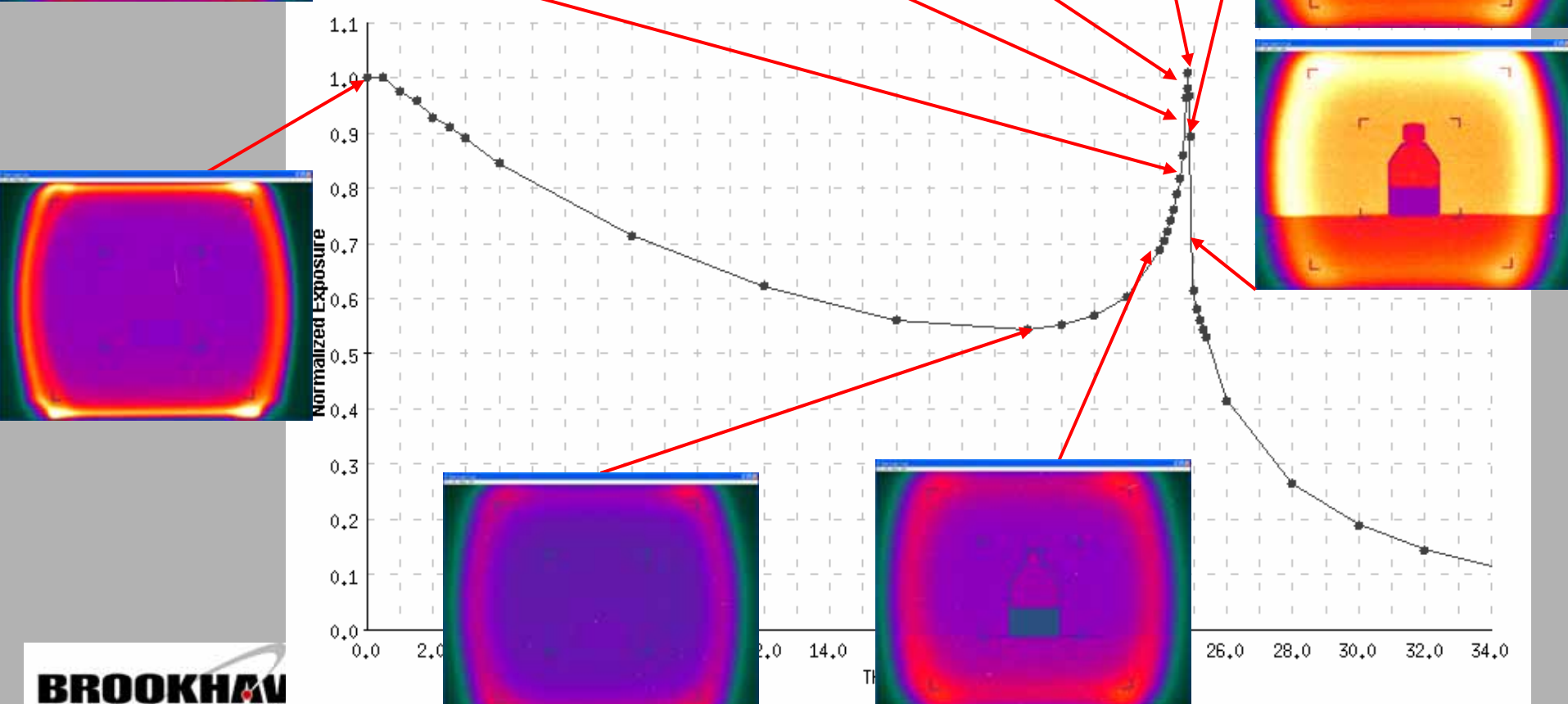
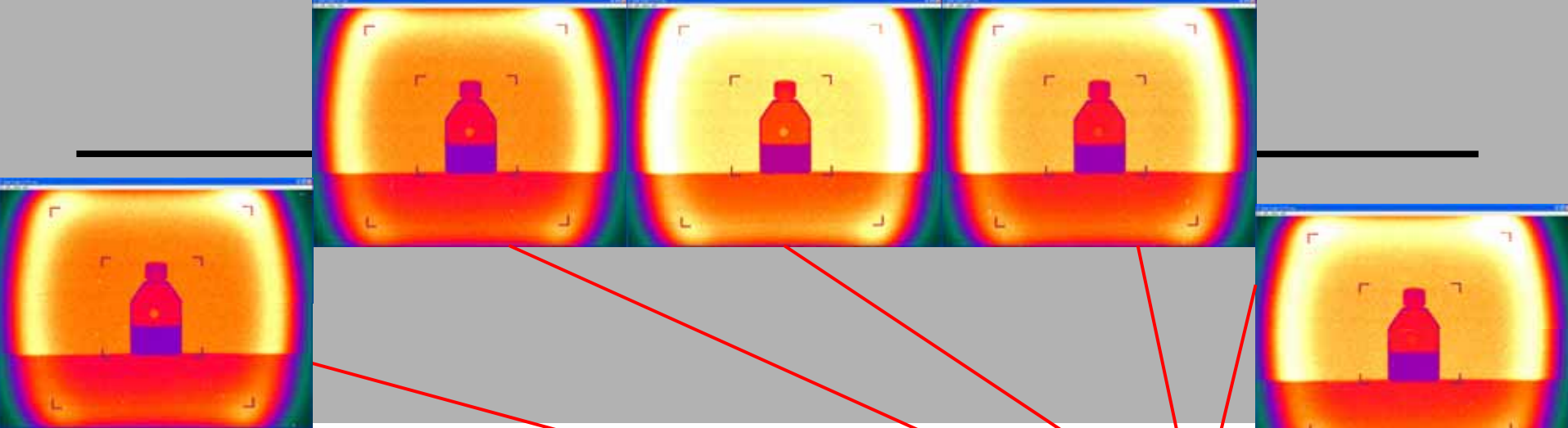
Charlie Pearson
Engineering design



Beam Profile Camera System
Side Elevation View







Outline

The NASA/BNL Space Radiation Program

Sampling of experiments

The NSRL Facility

Operations

Beam Characteristics

Uniform Beams

Dosimetry

Large Dynamic Range Camera Imaging System

Solar Particle Simulator

Summary

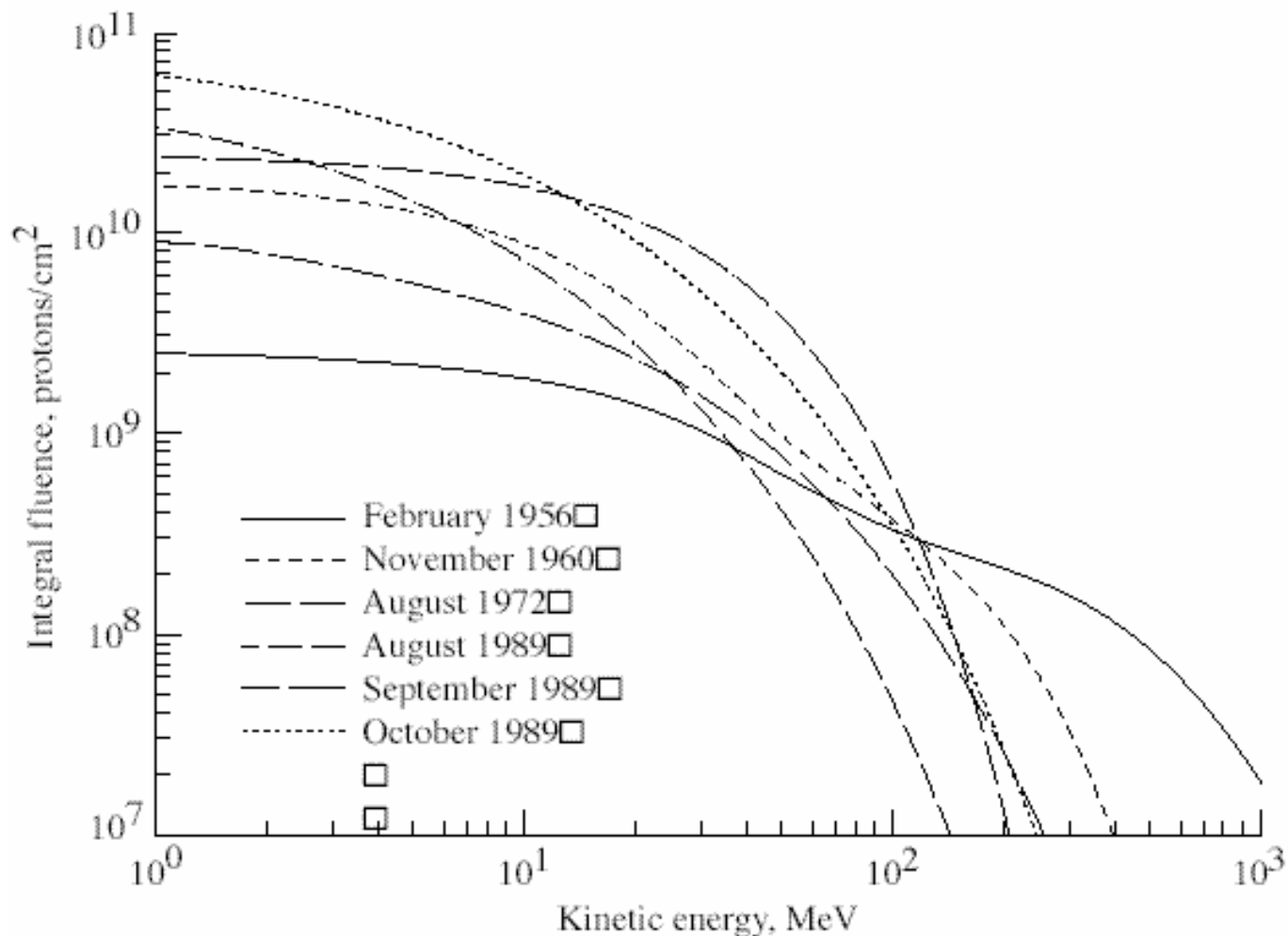
Solar Particle Simulator

The main motivation is to be able to reproduce energy spectra of the environment in space, particularly solar events.

NSRL Ground rules

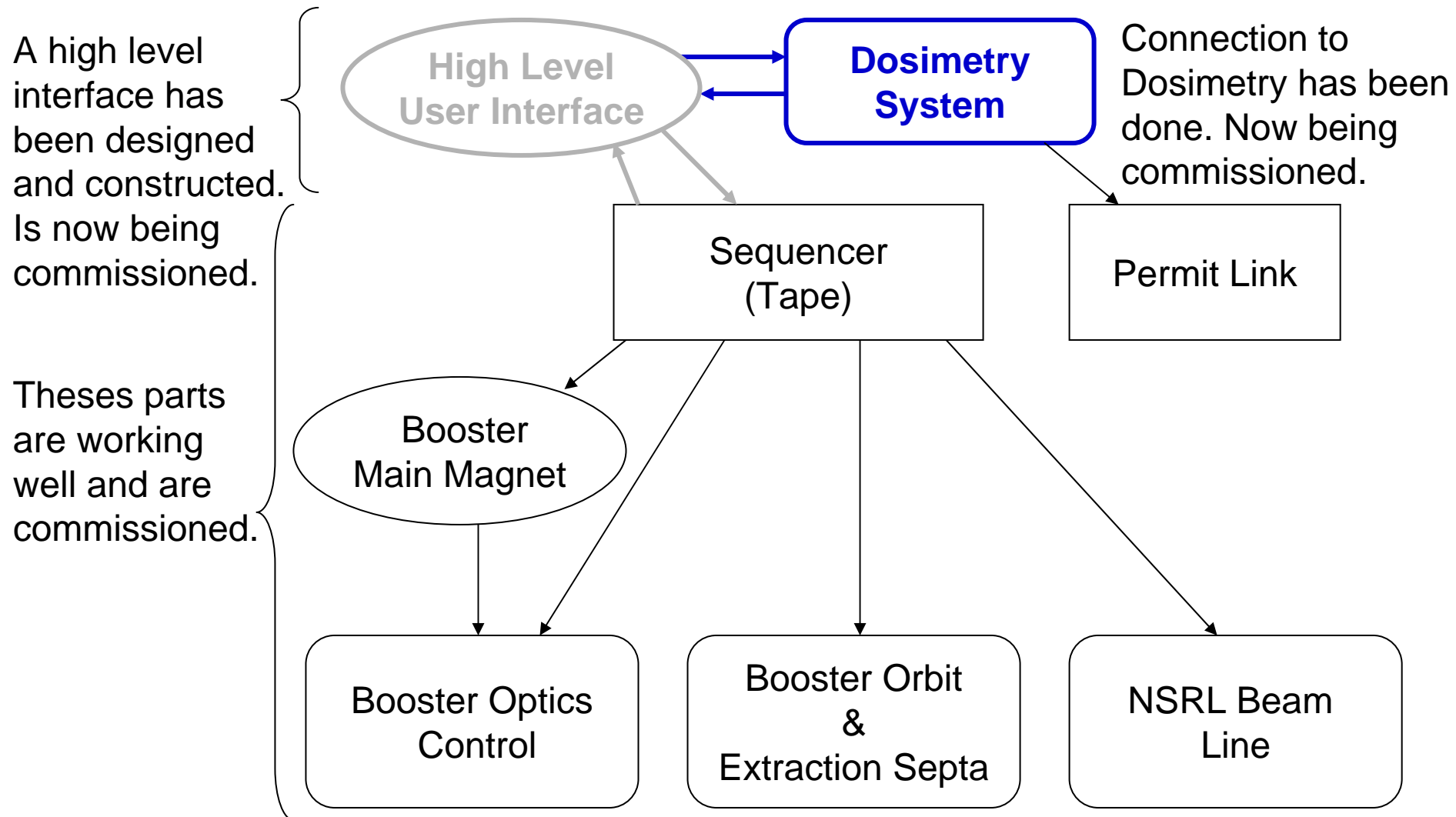
- *Need to be able to irradiate a single sample over entire energy spectrum within as little as a 1 hour period.*
- *Need a clean beam, as clean as any current NSRL experiment now receives.*
- *Need to know the energy of the beam as well as any current NSRL experiment.*
- *Need to know the dose as well as any current NSRL experiment.*

Solar Proton Events



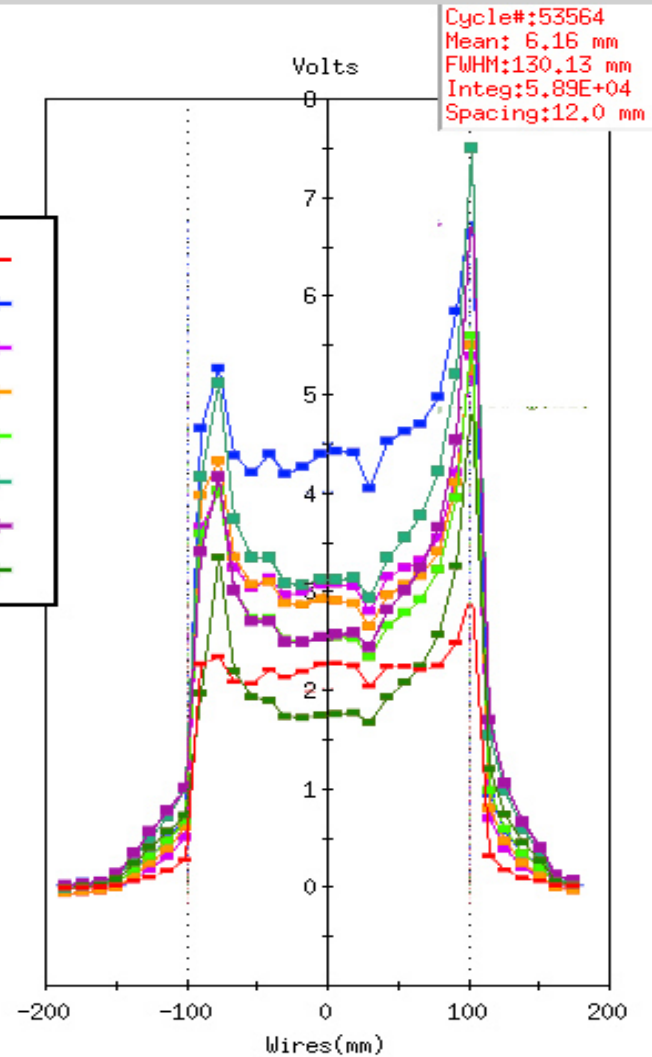
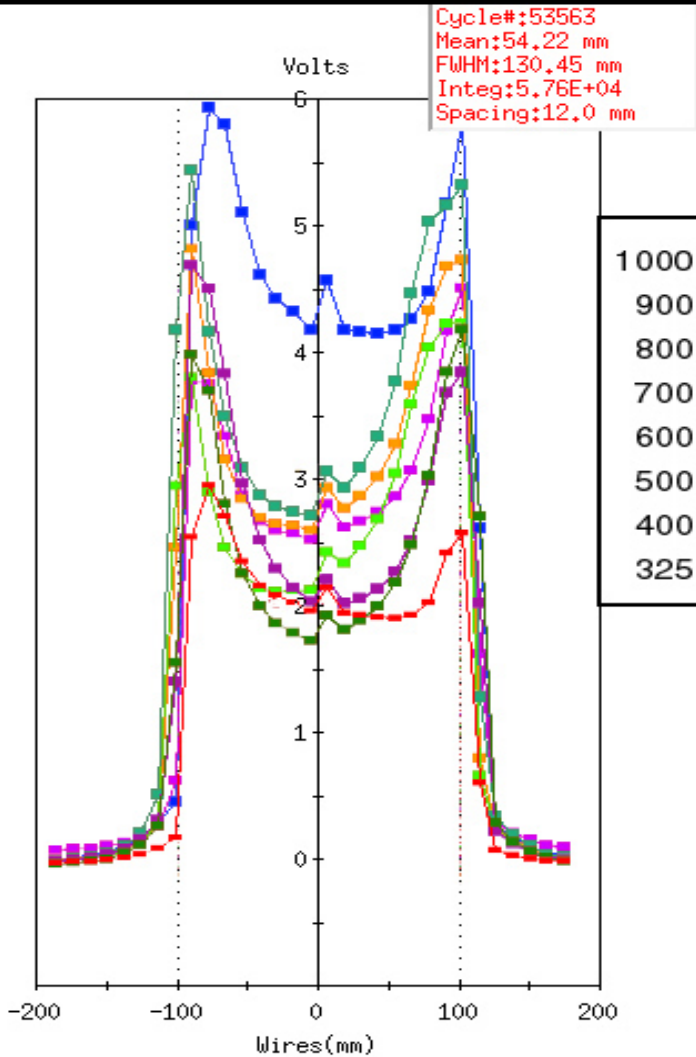
. Large solar proton event integral fluence spectra at 1 AU.

Actively Changing Energy at NSRL



Horizontal

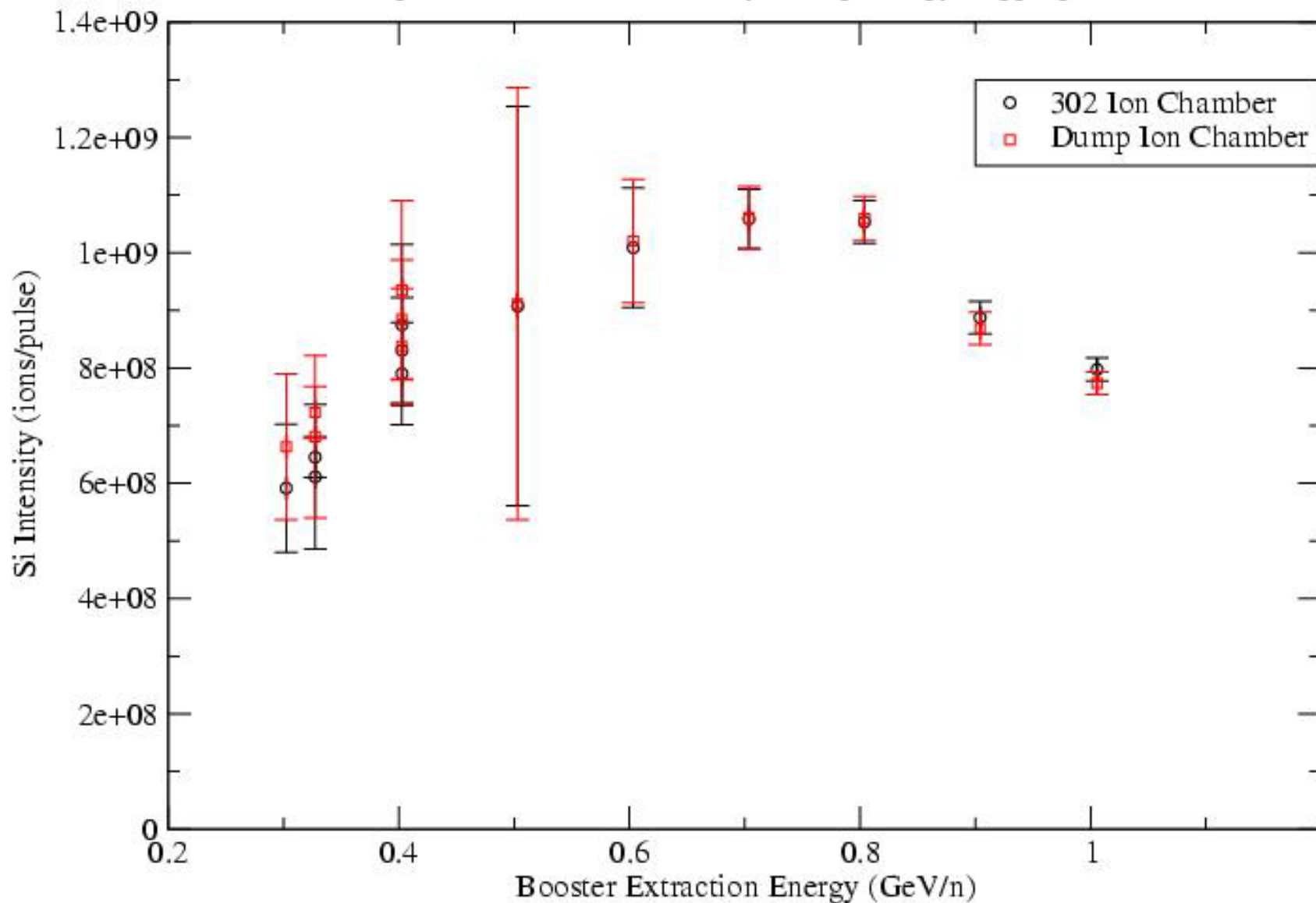
Vertical



BAF,R-SWIC302,H,CLD0

BAF,R-SWIC302,V,CLD0

Average Delivered Beam intensity during Energy stepping studies.



Outline

The NASA/BNL Space Radiation Program

Sampling of experiments

The NSRL Facility

Operations

Beam Characteristics

Uniform Beams

Dosimetry

Large Dynamic Range Camera Imaging System

Solar Particle Simulator

Summary

Current Capabilities

1. **Beam energies from 0.05 to 1 GeV/n with any ion that can be produced by a Tandem Van de Graff**
2. **Lowest intensities operated $\sim 10^2$ ions/cm²/cycle**
3. **Smallest beams around 1 cm, smaller possible (not uniform, Gaussian). Largest beams around 20x20 cm² uniform**
4. **Mixed field of protons and ions on a single target**
5. **1 – 3 msec pulsed beams**
6. **Fast Extracted Beams (1 to 3 200 nsec pulses in 5 usec, or one 4 usec pulse)**
7. **Solar particle simulator**
 1. **Large range of ion energies over single irradiation**
 2. **Fast energy change**

Future Capabilities

EBIS (electron beam ion source) will provide

- All ions up to U, including noble gases
- Higher intensities for current ion set
- Multiple mixed field ions (more than 2 ions species/irradiation)

Pulsed synchronized beams (with experimenter signals)

1. Synchronize beam with breathing, heart rate, EKG,
2. For low energy beams, allow for time of flight analysis

Supplemental

Slow Extraction Dynamics

Transverse fields in a normal sextupole are :

$$B_x(x, y) = -6B_s xy$$

$$B_y(x, y) = -3B_s(x^2 - y^2)$$

where,

$$B_s = -\frac{1}{6} \left(\frac{d^2 B_y}{dx^2} \right)_0$$

A particle with a magnetic rigidity $B\rho$ receives (thin lens) kicks by a sextupole of length L ,

$$\Delta x' = \frac{1}{2} \frac{L}{B\rho} \left(\frac{d^2 B_y}{dx^2} \right)_0 (x^2 - y^2)$$

and,

$$\Delta y' = -\frac{L}{B\rho} \left(\frac{d^2 B_y}{dx^2} \right)_0 (xy)$$

define the normalized sextupole strength as

$$S = \frac{1}{2} \beta^{3/2} \frac{L}{B\rho} \left(\frac{d^2 B_y}{dx^2} \right)_0$$

The Kobayashi Hamiltonian

$$H = \frac{\varepsilon}{2} (X^2 + X'^2) + \frac{S}{4} (3XX'^2 - X^3)$$

where

$$X \equiv \frac{x}{\sqrt{\beta_x}} \quad \text{and} \quad X' \equiv \frac{\alpha_x}{\sqrt{\beta_x}} x + \sqrt{\beta_x} x'$$

are the normalized phase space coordinates,

and,

$$\varepsilon = 6\pi\delta Q = 6\pi(Q_{particle} - Q_{resonance})$$

The first term in H describes particle motion in the linear unperturbed lattice ($S = 0$). The trajectories are circles with radius $\sqrt{2H / \varepsilon}$ in normalized phase space.

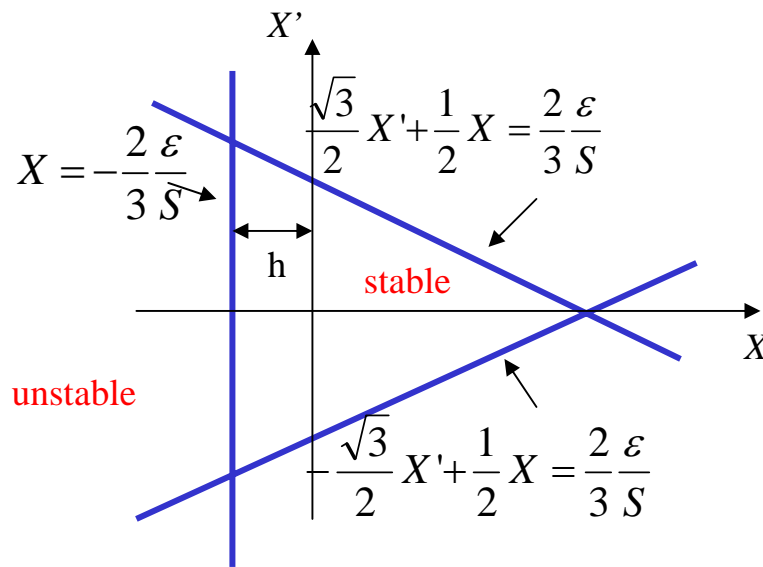
The second term (perturbative term) distorts the circle.

Slow Extraction Dynamics

when H has the value $\left[(2\varepsilon/3)^3/S^2\right]$ it factors into 3 terms,

$$\left(\frac{S}{4}X + \frac{\varepsilon}{6}\right)\left(\sqrt{3}X' + X - \frac{4\varepsilon}{3S}\right)\left(\sqrt{3}X' - X + \frac{4\varepsilon}{3S}\right) = 0$$

The 3 lines define the boundaries between stable and unstable regions of phase space. The size of the stable region is determined by the ratio ε/S .



$$h = \frac{2\varepsilon}{3S} = \frac{4\pi}{S} \delta Q$$

The area of the stable region is

$$A = 3\sqrt{3}h^2 = \frac{48\sqrt{3}\pi}{S^2} (\delta Q^2) \pi$$

The area within a particular particles linear unperturbed motion is called the single particle emittance.

$$E = a^2 \pi, \text{ where } a^2 = X_0^2 + X_0'^2$$

Particle motion remains stable as long as the particle motion lies within the stable triangle.

$$E_{stable} \leq \frac{48\sqrt{3}\pi}{S^2} (\delta Q^2) \pi$$

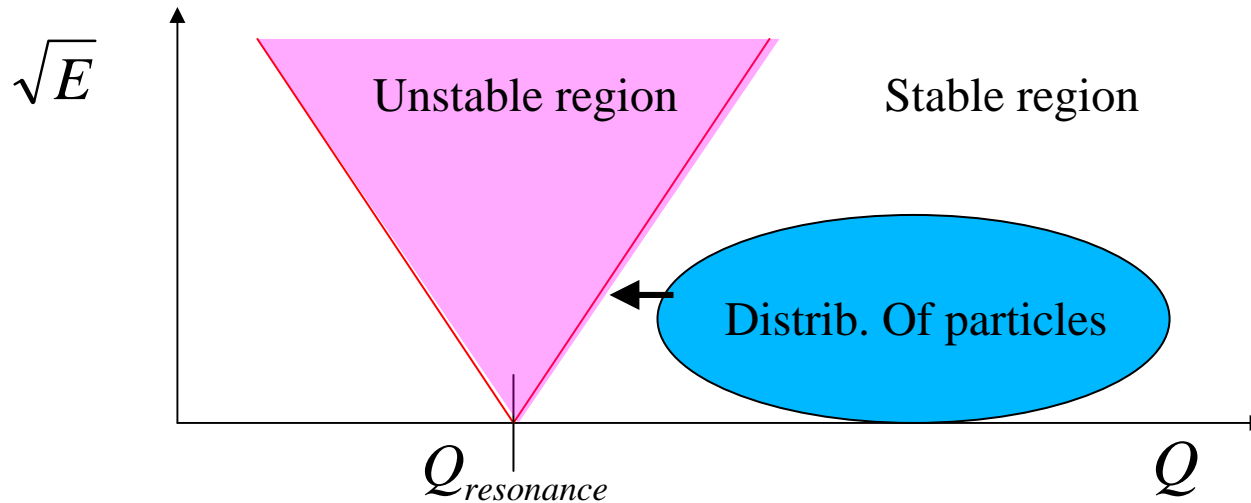
This criteria for stability can be rewritten in terms of absolute betatron tune, with

$$\delta Q = Q_{particle} - Q_{resonance}$$

$$Q_{resonance} - \sqrt{\frac{1}{48\sqrt{3}\pi}} S \sqrt{\frac{E}{\pi}} < Q_{particle} < Q_{resonance} + \sqrt{\frac{1}{48\sqrt{3}\pi}} S \sqrt{\frac{E}{\pi}}$$

This can be shown graphically by plotting the action variable, \sqrt{E} , as a function of betatron tune.

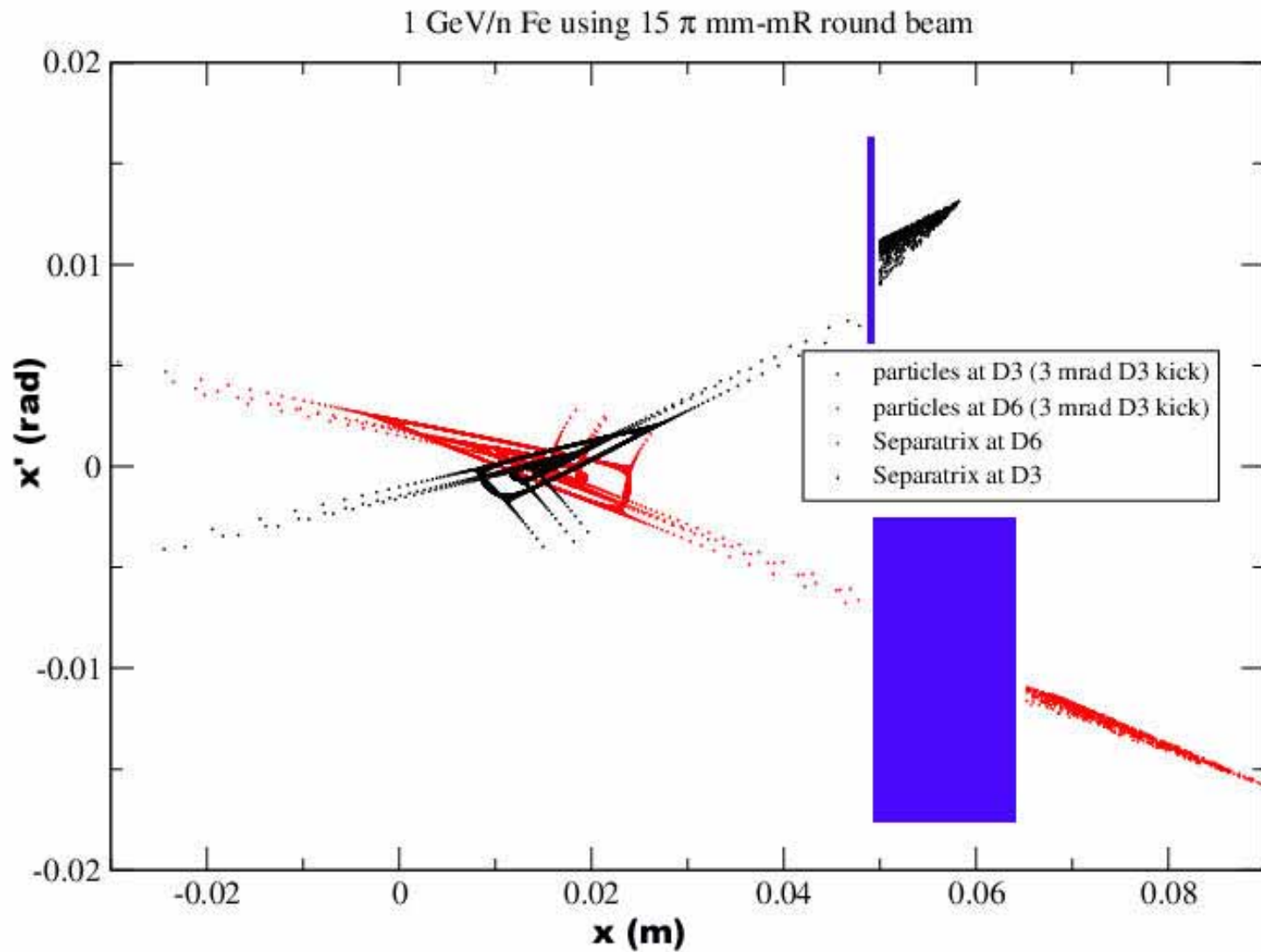
Slow Extraction Dynamics



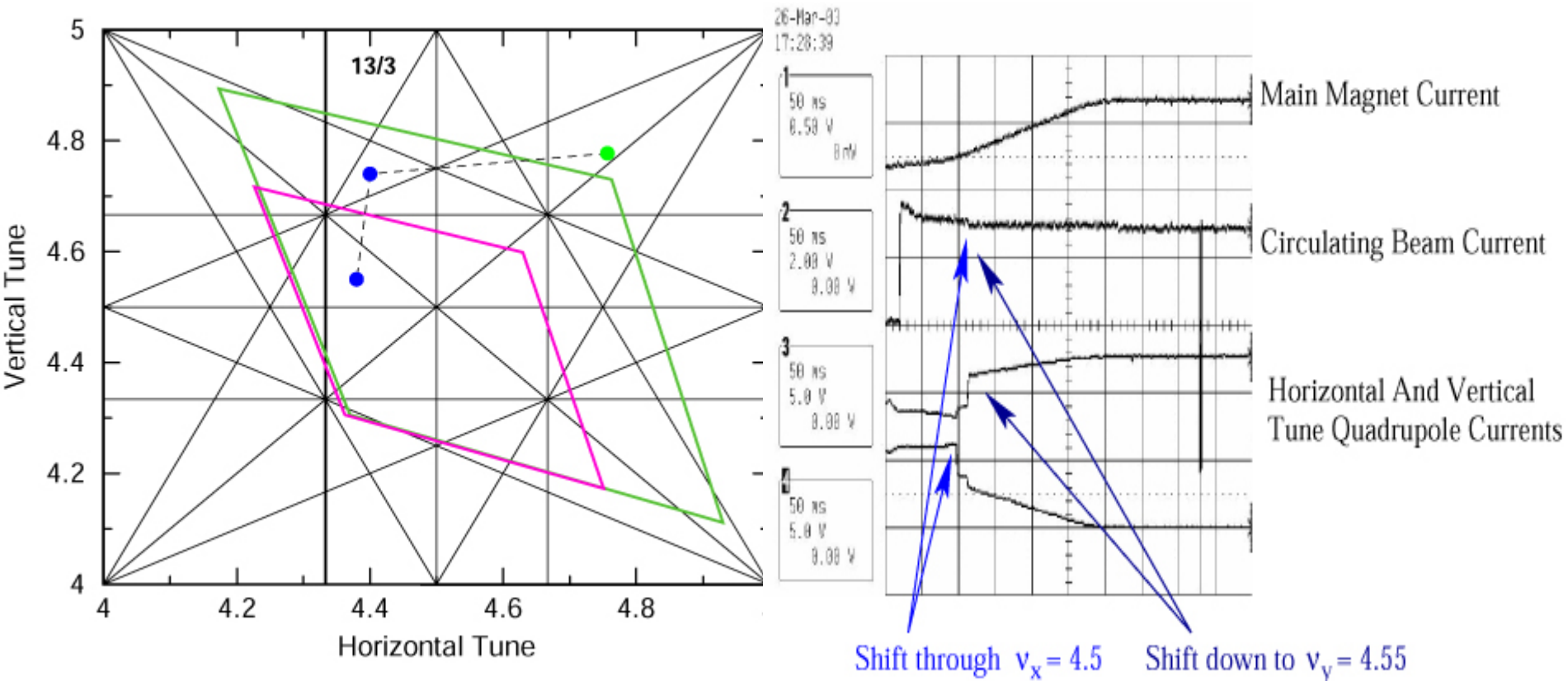
Extraction Methods:

1. Widen stopband by increasing S
2. Move particles into resonance by changing betatron tune (AGS).
3. Increase particle amplitudes until it encounters the unstable region (rf knockout method).

Booster Resonant/Slow Extraction



Injection, Acceleration, and Tune Space Manipulations



Scope Selector :

LeCroy 9414, 9354, LC584AL"

Comment:

STOP

05/30/03 02:49 PM

Scope settings

time/div 2 ms

delay 0.0E+0
<0:ns , >0:%

memory 250 k

Bunch Ampl Factor 12000

RF freq. (MHz) 3.6872400

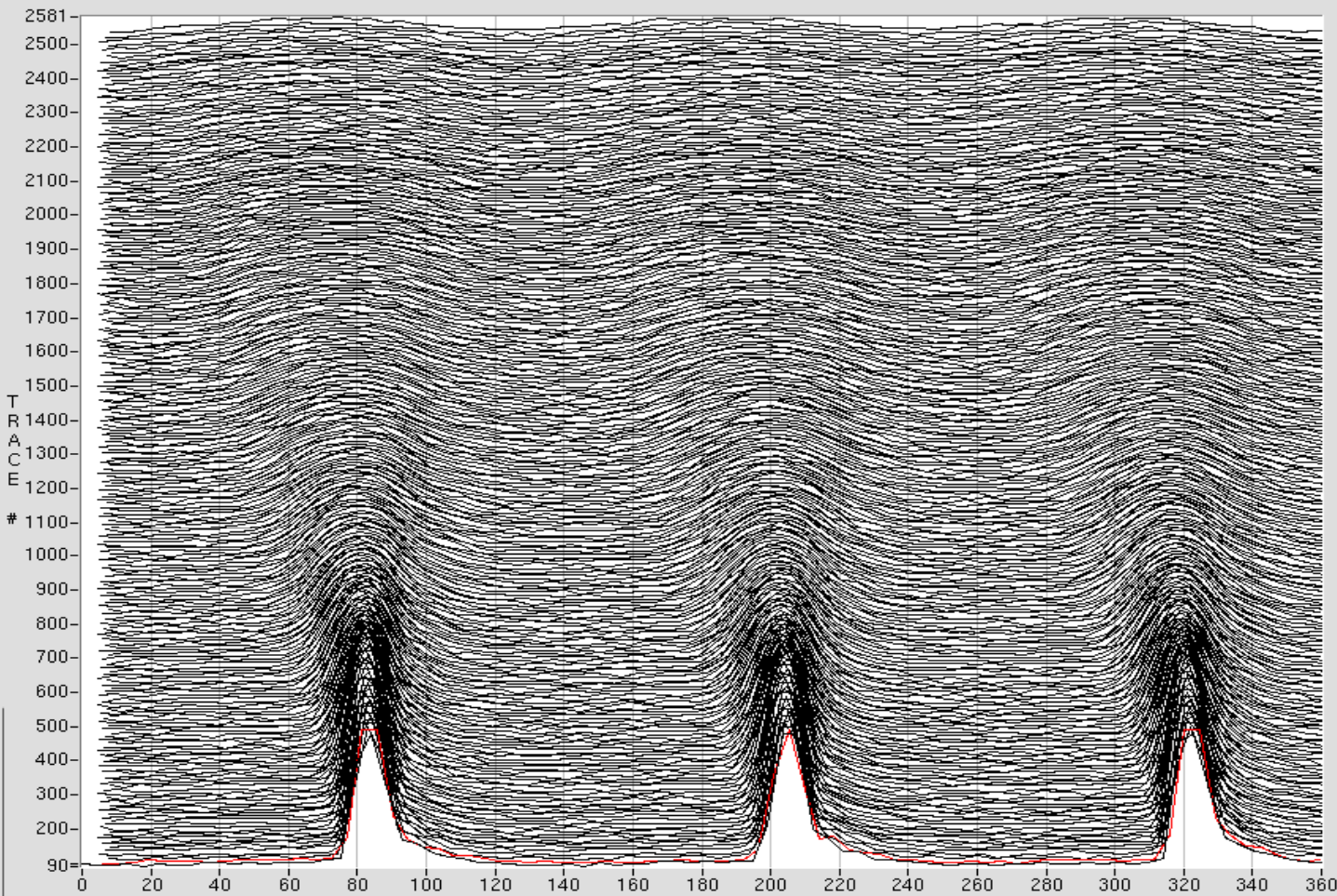
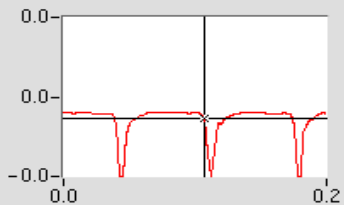
df/dt khz/ms 10.00

harmonic # 3

turns per trace 1

Sampling time (ns) 10.00

trace # in icon 10



Cursor 0.086 -0.01

GPIB address 10:

TRIGGER SETUP

Local Selected

TRIGGER LEVEL

2.00

/home/cfsb/mcr/labview/data

CH4 data entry TEMP.XLS
volt/div 10 mV offset (mV) 0

DISPLAY CH1 1 CHANNEL MODE
Settings "ABOVE"

CH 1 data entry data1
volt/div 500 mV ch1 offset (mV) 0

GEOMETRIC PHASE (deg)

RE_DISPLAY NEGATED SIGNAL

from Memory

Display every "X" Trace 10

Energy Measurement in an Accelerator

Energy

$$\beta = (2 \pi R_{\text{booster}} f_{\text{rf}}) / (h c)$$

$$\gamma = 1/(1-\beta^2)^{1/2}$$

$$\text{K.E.} = (\gamma - 1)mc^2$$

Uncertainty

$$d\beta^2 = (2\pi f_{\text{rf}} / hc)^2 dR^2 + (2\pi R/hc)^2 df_{\text{rf}}^2$$

$$d(\text{K.E.}) = mc^2 \beta (1-\beta^2)^{-3/2} d\beta$$

Symbols:

h = harmonic #

c = 2.99792458x10⁸ m/s

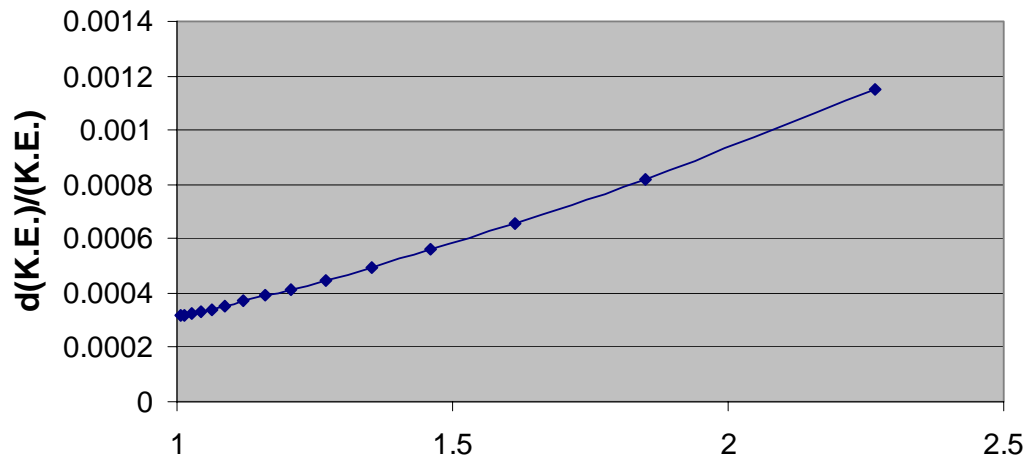
R_{booster} = R = 32.113 m

f_{rf} = accelerating frequency

Values used in figure:

dR = 5 mm df_{rf} = 10 Hz h = 3

Fractional Uncertainty in K.E. vs γ



Spill and Ripple

Time Structure Formalism

definitions:

Q = horizontal betatron tune

ξ = horizontal chromaticity = $\frac{dQ/Q}{dp/p}$

I_m = current in the Main Dipoles and Quadrupoles

N = number of particles ($\frac{dN}{dQ}$ represents the particle distribution in tune space)

T = period over which particles are extracted

Low frequency duty factor:

$$D_f = \frac{[f_T S(t) dt]^2}{T f_T [S(t)]^2 dt} \Rightarrow \left(\frac{f_{av}}{f_{rms}}\right)^2 \text{ in general}$$

where

$$S(t) = \frac{dN}{dt} = \frac{dN}{dQ} \frac{dQ}{dt}$$

if there is no ripple,

$$S(t) = \frac{dN}{dQ} \dot{Q}_0$$

where \dot{Q}_0 is the rate at which particles move into resonance.

$$\dot{Q}_0 = \frac{Q\xi}{I_m} \frac{dI_m}{dt}$$

Spill and Ripple

If there is ripple on the magnet power supplies;

$$S(t) = \frac{dN}{dQ}(\dot{Q}_0 + \dot{Q}_v)$$

where \dot{Q}_v is the variations in the rate at which particles move into resonance.

$$\dot{Q}_v = \frac{Q\xi}{I_m L_m} \sum_h V_h$$

L_m is the total inductance of the main dipoles and quads and V_h is the sum of the 60 Hz harmonics amplitudes (in volts).

Reducing Time structure using RF Phase Displacement

$$S(t) = \frac{dN}{dQ} \dot{Q}_0 \left(1 + \frac{\dot{Q}_v}{\dot{Q}_0}\right)$$

For 1 particular frequency we can write the duty factor as

$$D_f = \frac{1}{1 + \frac{1}{2}\left(\frac{\dot{Q}_v}{\dot{Q}_0}\right)^2} = \frac{1}{1 + \frac{1}{2}\left(\frac{\omega\delta Q}{v_0}\right)^2}$$

where

ω = frequency

δQ = relative ampl. of that freq. in tune space

v_0 = speed that beam crosses resonance

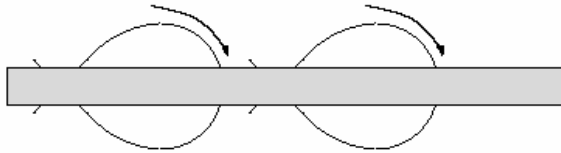
$$v_0 = \frac{\Delta p}{p} \frac{1}{T}$$

Spill and Ripple

D_f is increased by

1. decreasing δQ
2. increasing v_0

One way to increase v_0 is to increase $\frac{\Delta p}{p}$. To further increase it we use RF phase displacement, using a high frequency RF cavity. In this case RF buckets are centered on the resonance.



The buckets are empty and beam is forced between them.

Now,

$$D_f = \frac{1}{1 + \frac{RB\rho T}{V\frac{\Delta p}{p}}(\omega\delta Q)^2}$$

Without RF phase displacement, a 100 % modulated spill has $D_f = 0.67$. In this case,

$$\omega\delta Q \geq \frac{\Delta p}{p} \frac{1}{T}$$

Spill and Ripple

